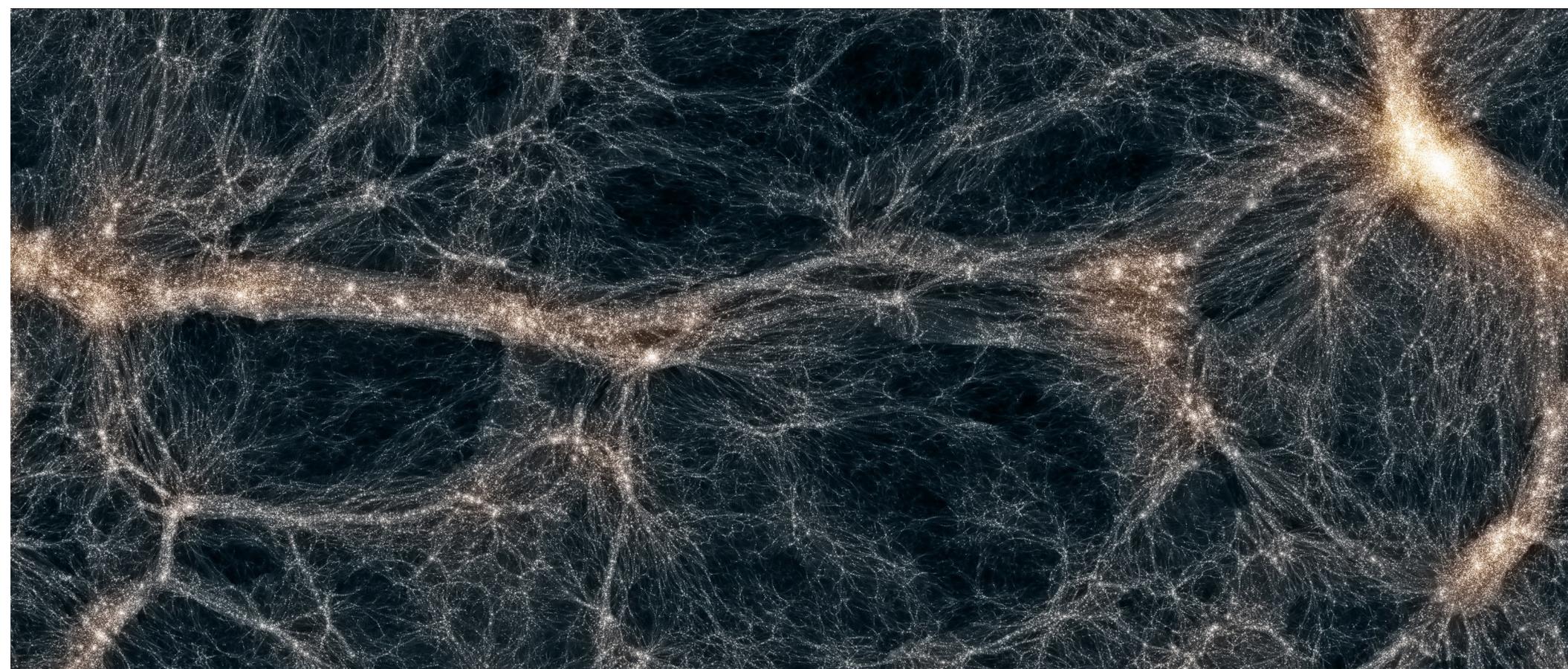


A Universe of ChaNGa Applications



Thomas Quinn
University of Washington



Ben Keller
Isaac Backus
Michael Tremmel
Joachim Stadel
James Wadsley
Charlotte Christensen
Spencer Wallace



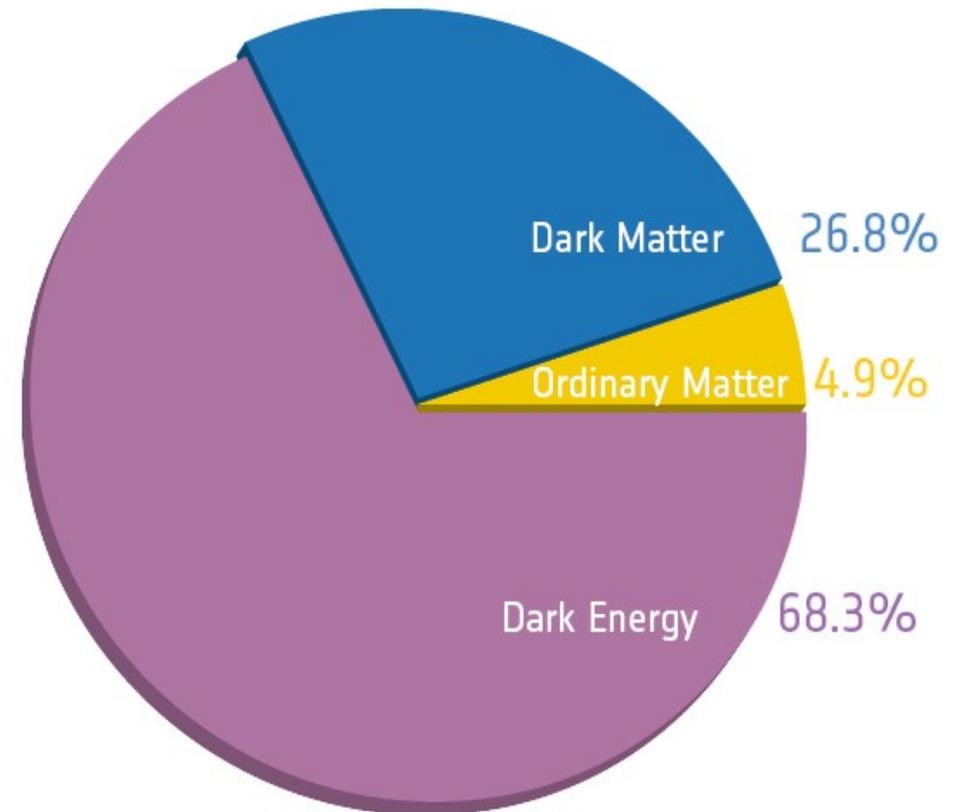
Filippo Gioachin
Pritish Jetley
Lukasz Wesolowski
Gengbin Zheng
Harshitha Menon
Orion Lawlor
Michael Robson
Jaemin Choi
Ronak Buch

Others:

Jianquo Liu, Purdue
Tim Haines, UW-Madison
Phil Chang, UW-Milwaukee

Fundamental Problem: Dark Matter and Energy: What is it?

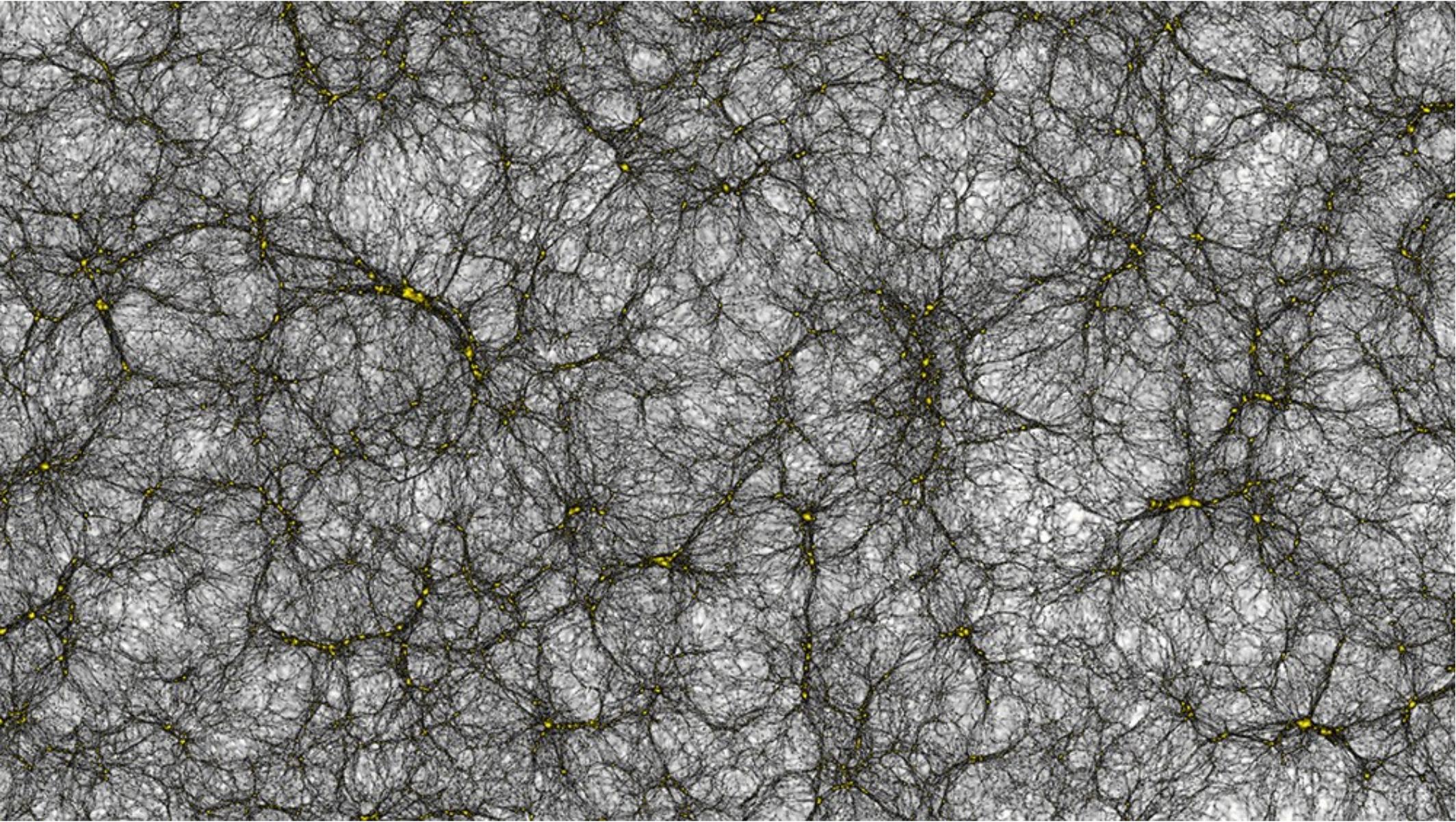
- Not baryons
- Gravitates!
- **Simulations** show:
not known neutrinos
- Candidates:
 - Sterile Neutrinos
 - Axions
 - Lightest SUSY Particle (LSP)

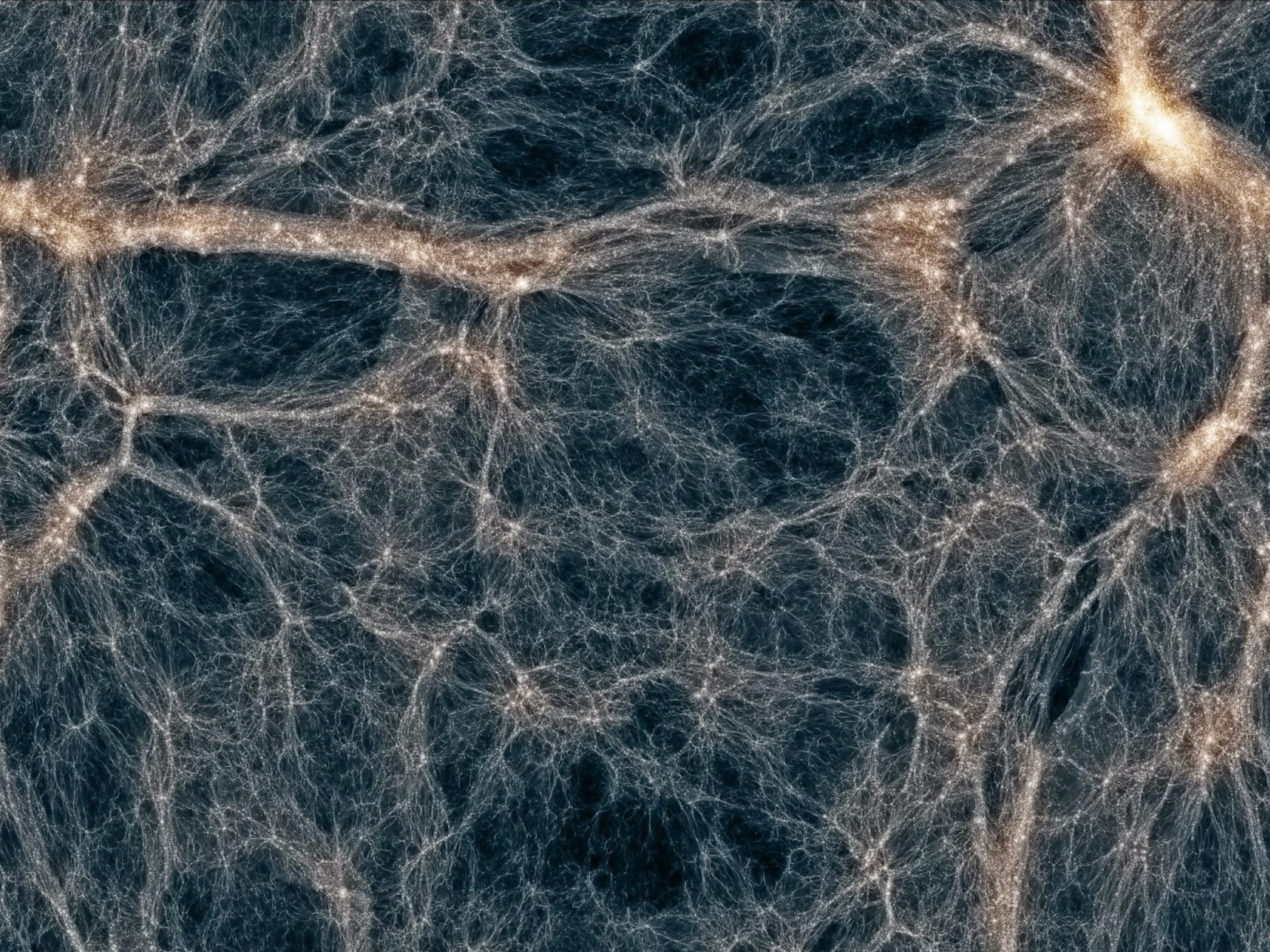


Modeling Dark Matter

- Physics is simple: Newton's Laws
- Computation is challenging: Naively order N^2
- Large spacial dynamic range: > 100 Mpc to < 1 kpc
 - Hierarchical, adaptive gravity solver is needed
- Large temporal dynamic range: 10 Gyr to < 1 Myr
 - Multiple timestep algorithm is needed
- Gravity is a long range force
 - Hierarchical information needs to go across processor domains

Gigaparsecs: the Cosmic Web





CHANGA

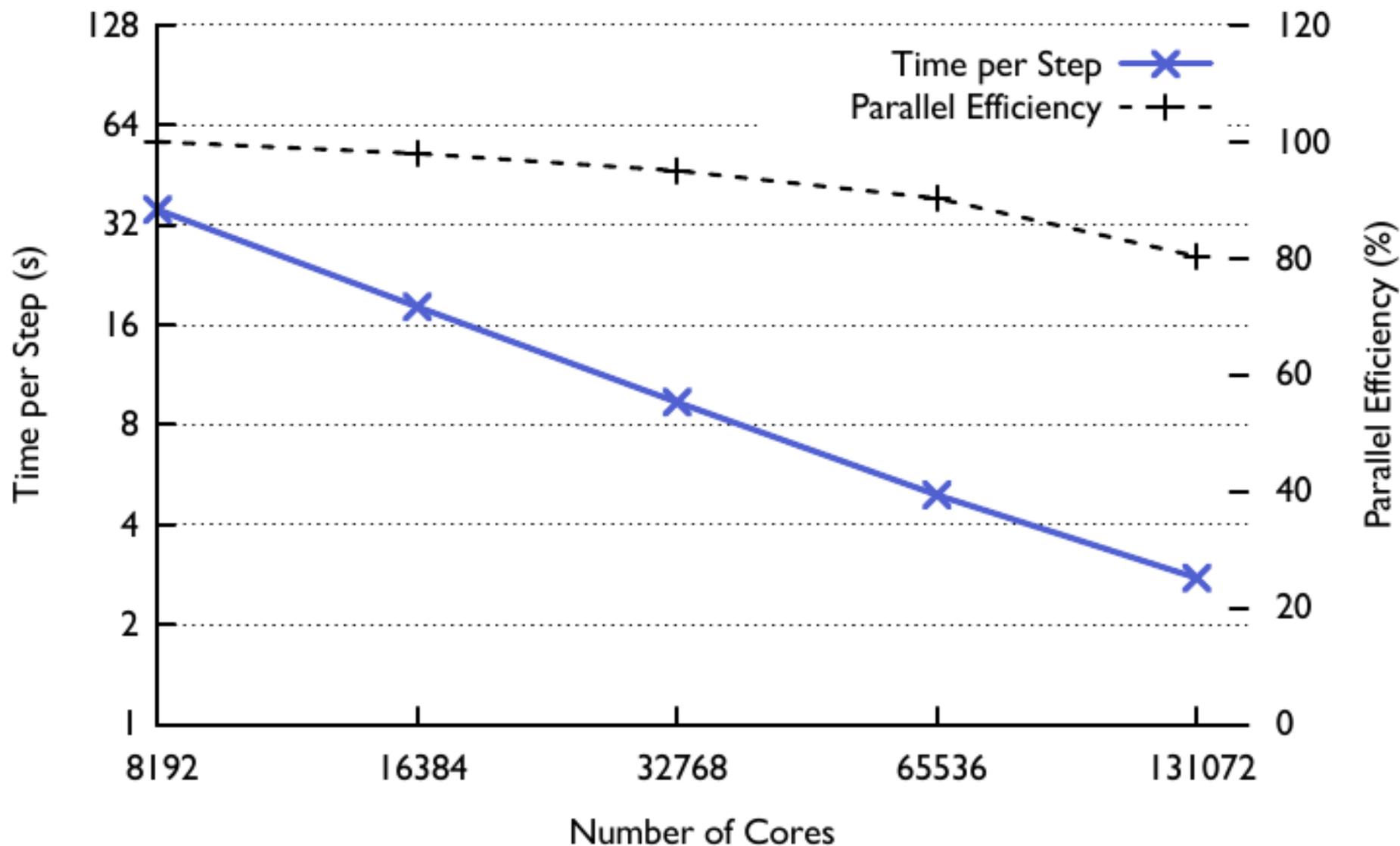


UNLEASHED

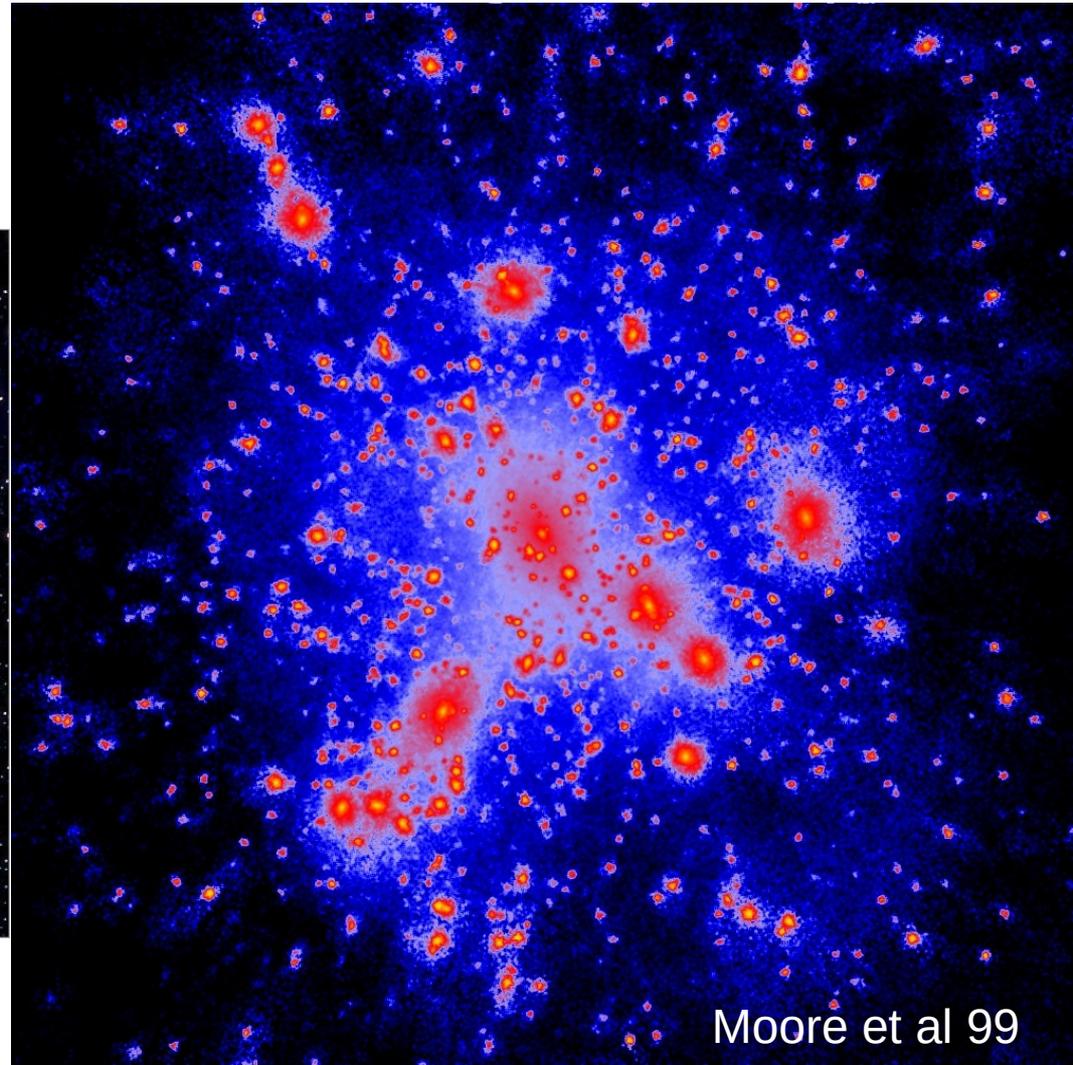
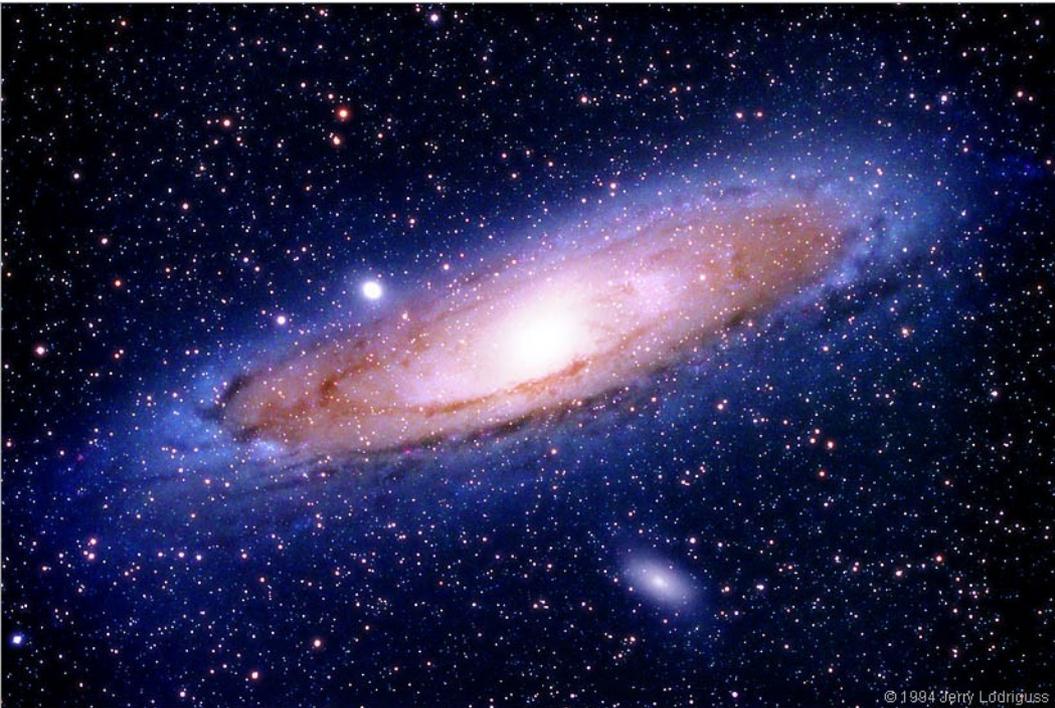
Charm Nbody GrAavity solver

- Massively parallel Tree-code
- Framework for Gravity and Neighbor Finding
 - K-nearest
 - Fixed radius
- Large number of Astrophysics models

Speedups for 2 billion clustered particles



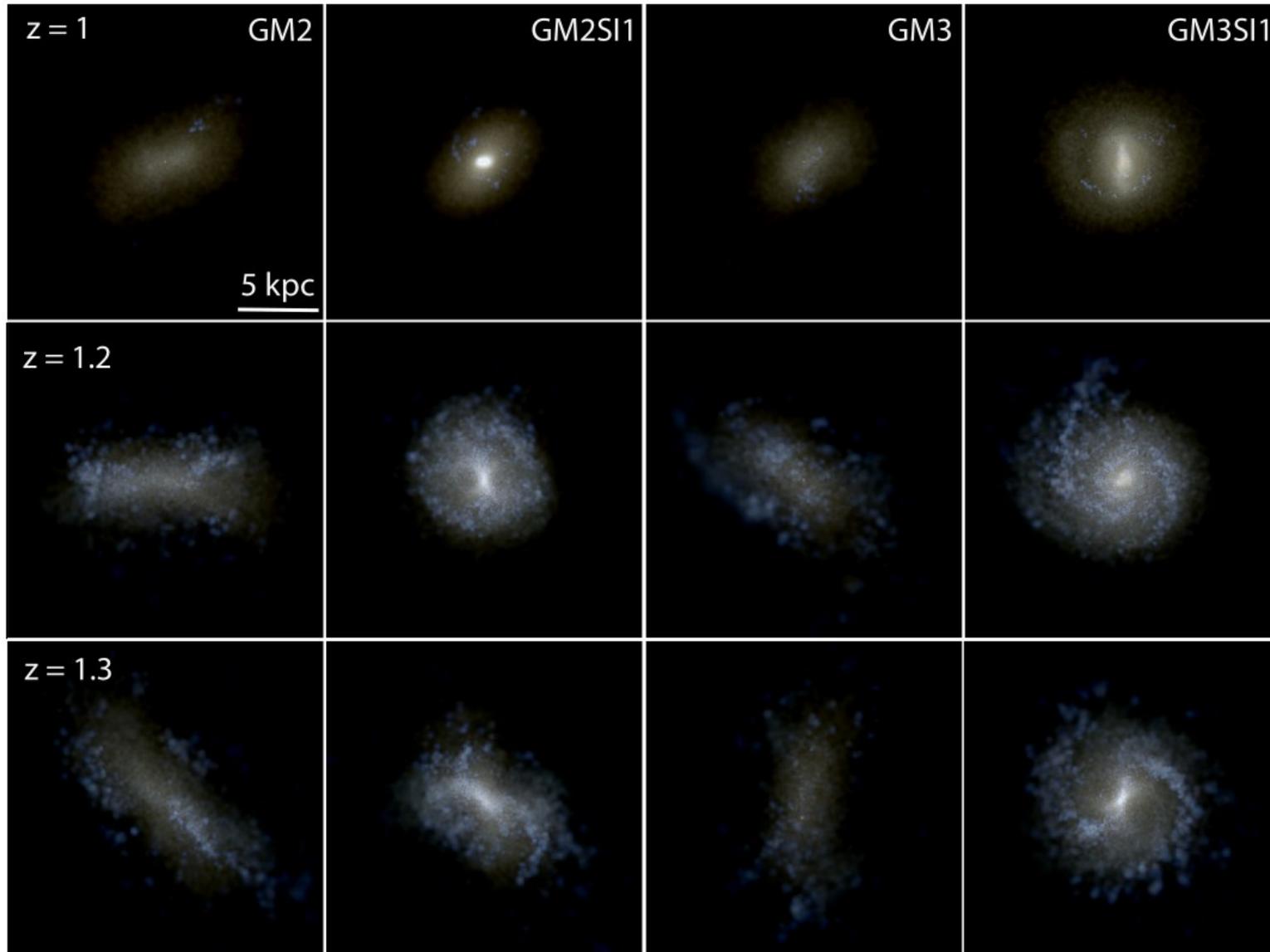
Light vs. Matter



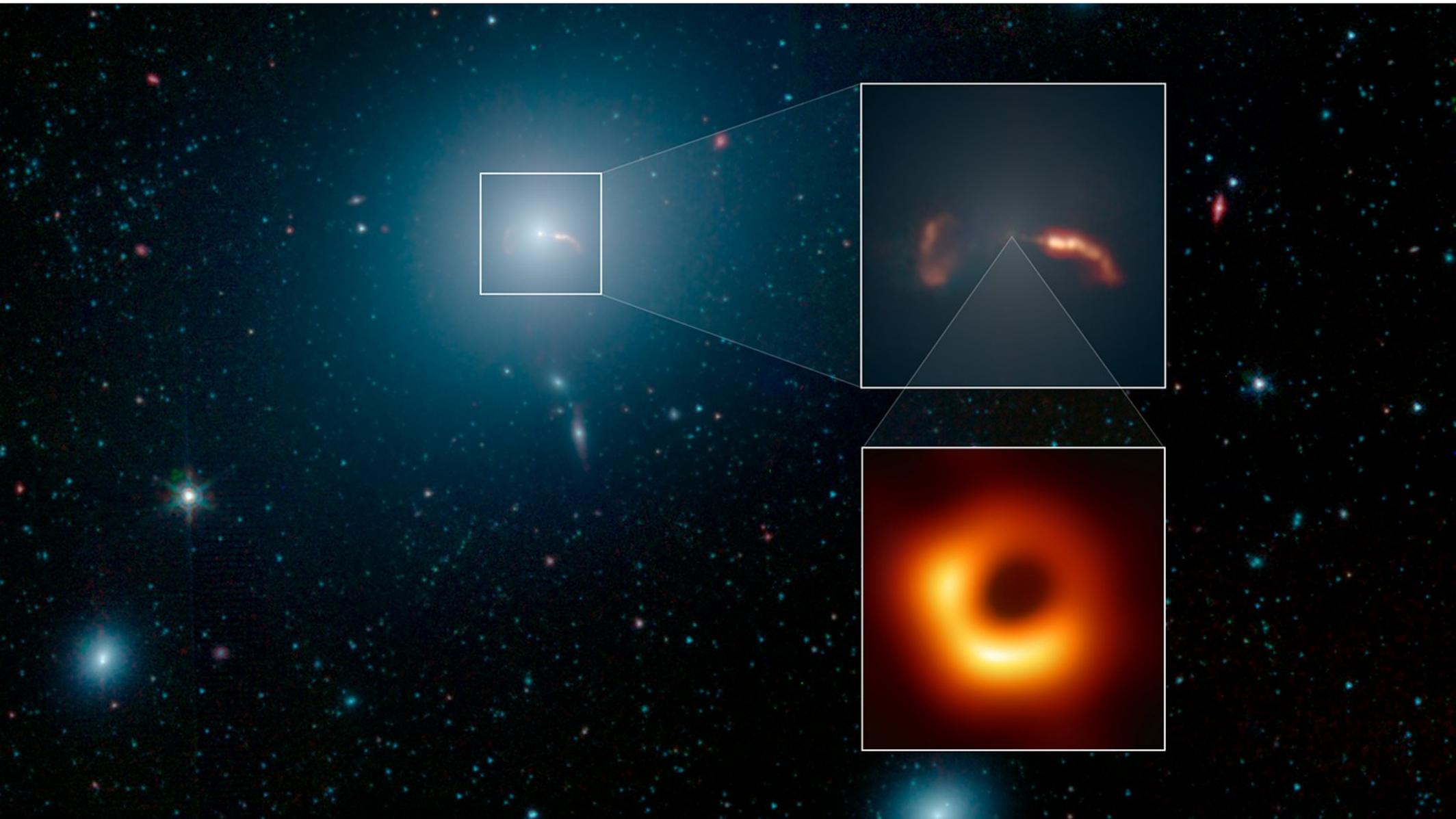
Smooth Particle Hydrodynamics

- Making testable predictions needs
Gastrophysics
 - High Mach number
 - Large density contrasts
- Gridless, Lagrangian method
- Galilean invariant
- Monte-Carlo Method for solving Navier-Stokes equation.
- Natural extension of particle method for gravity.

Testing Dark Matter Models



Black Holes!



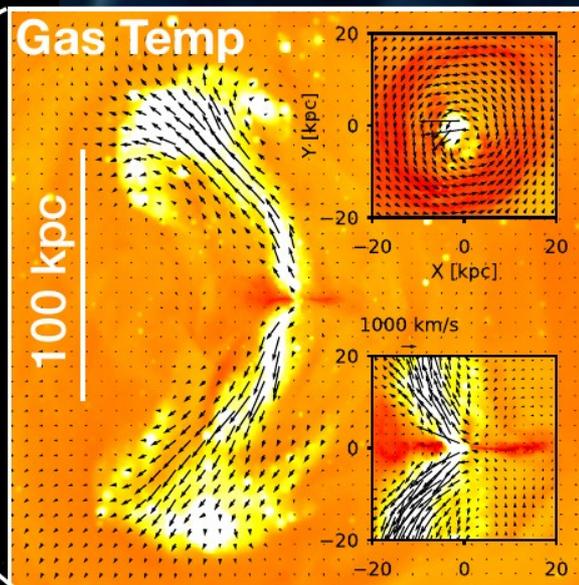
The ROMULUS Simulations

Certified organic, free-range, locally grown supermassive black holes

- ✓ Early Seeding in low mass halos
- ✓ Self-consistent and physically motivated dynamics, growth, and feedback
- ✓ Naturally produces large-scale outflows
- ✓ **No unnecessary additives or assumptions**

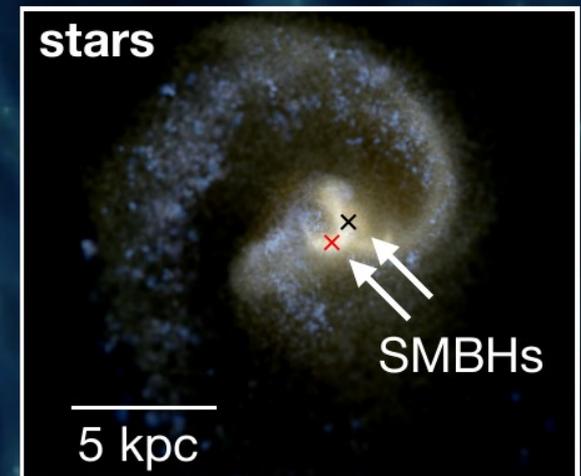
ROMULUSC

$10^{14} M_{\text{sun}}$ Galaxy Cluster
Tremmel+ submitted
(stars, uvj colors)



ROMULUS25

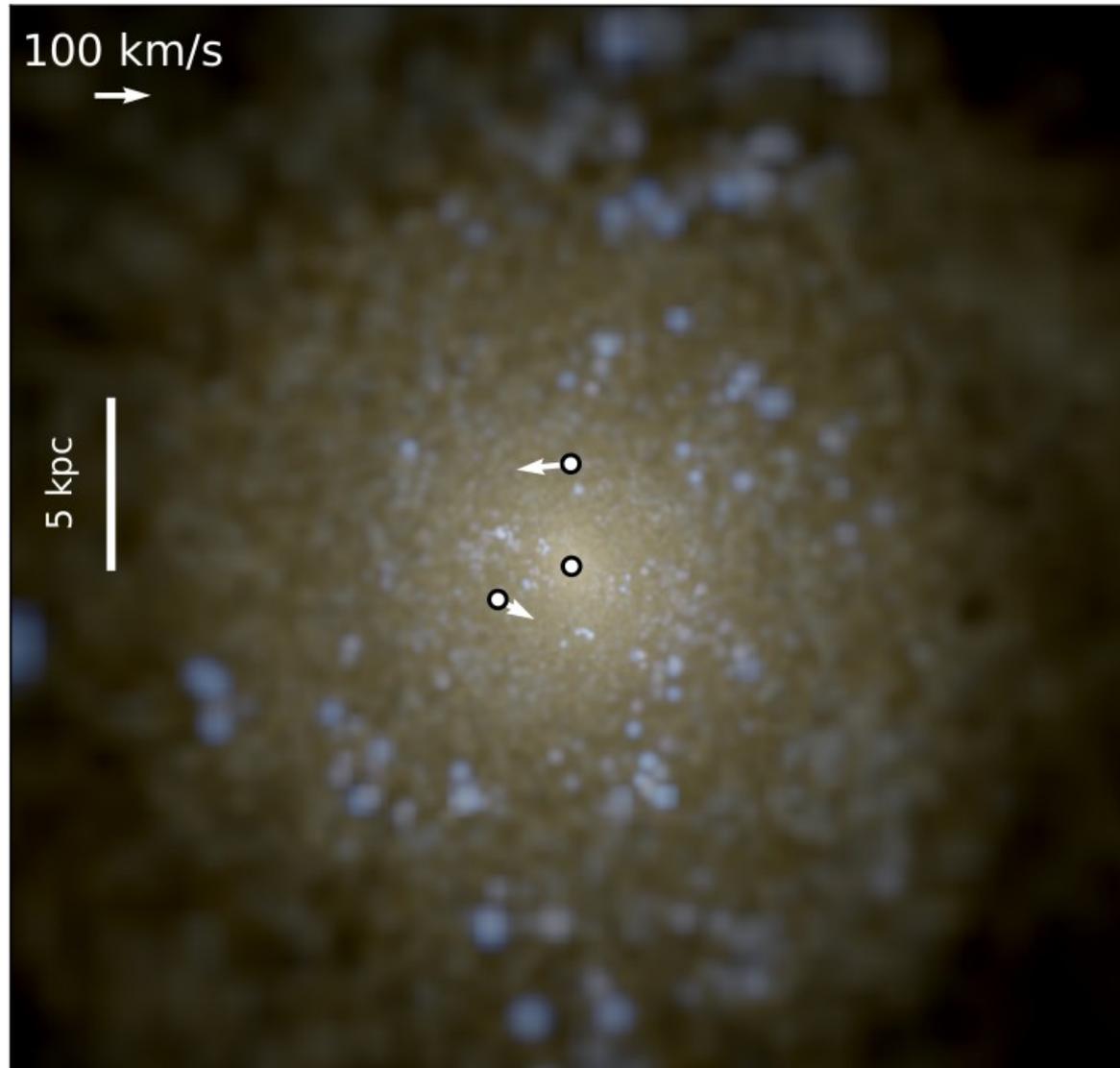
25 Mpc Volume
Tremmel+ 2017
(gas temp)



Resolution:
250 pc (grav)
50 pc (hydro)
 $\sim 1e5 M_{\text{sun}}$

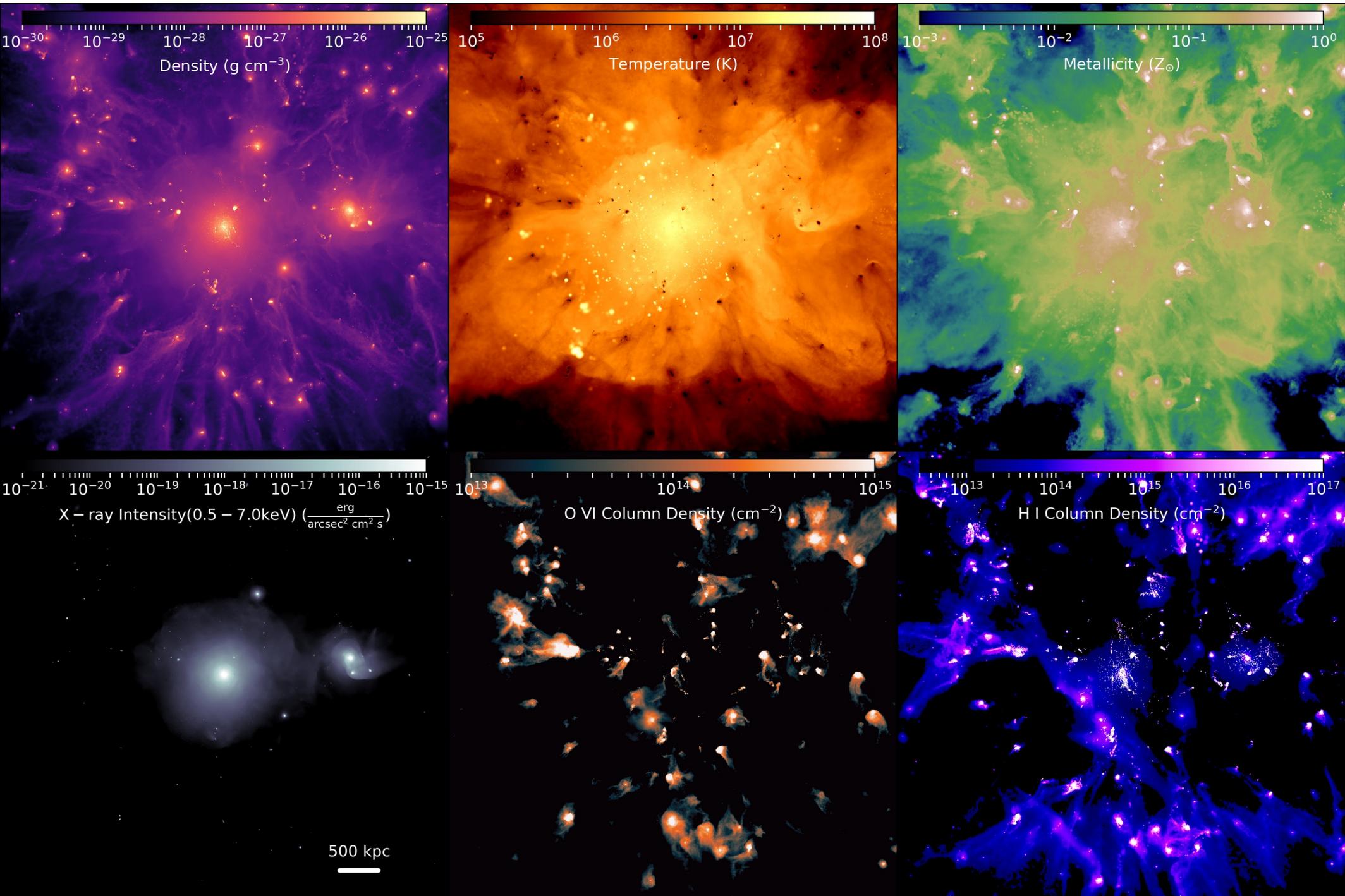
CHANGA

Wandering Black Holes?

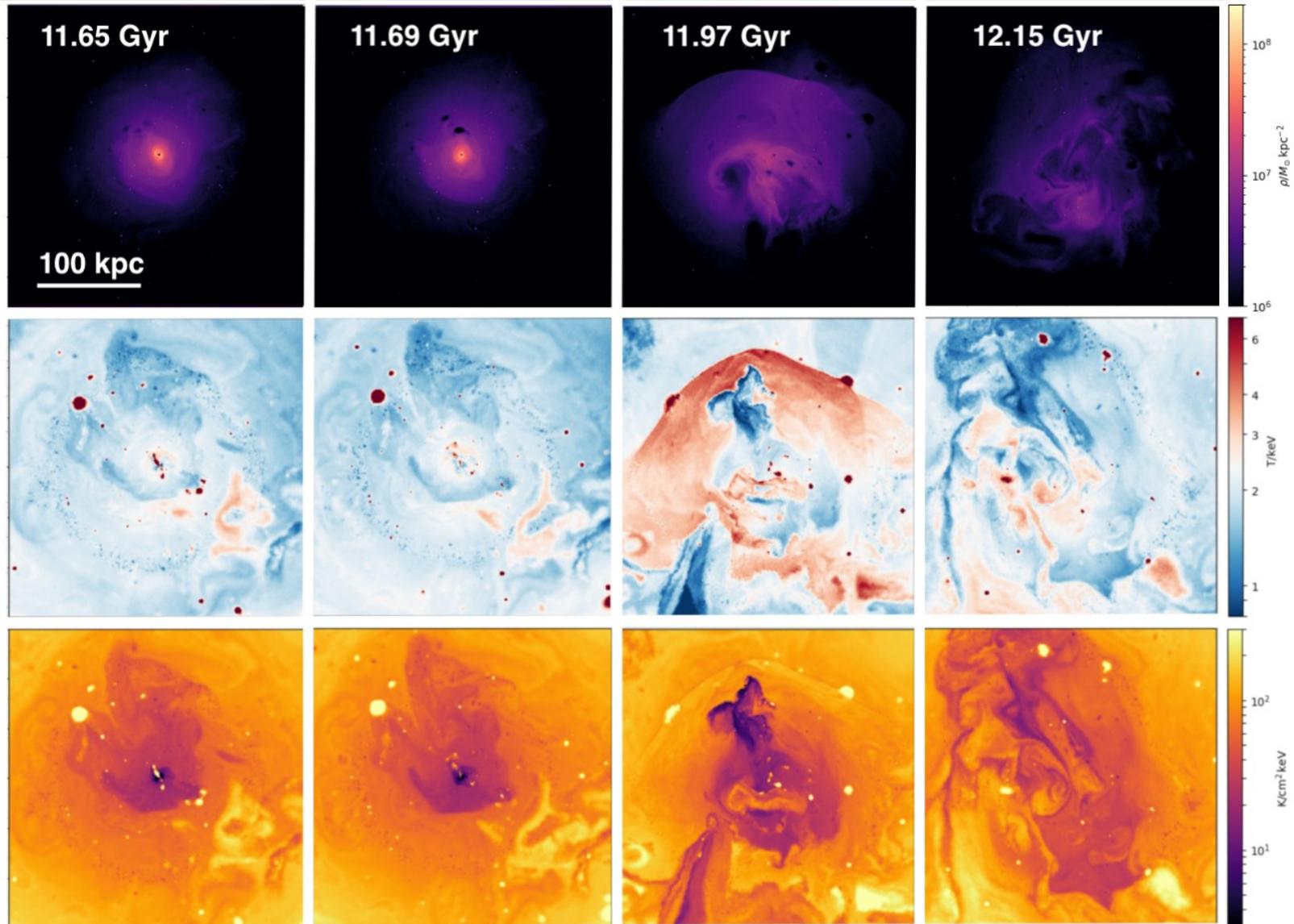




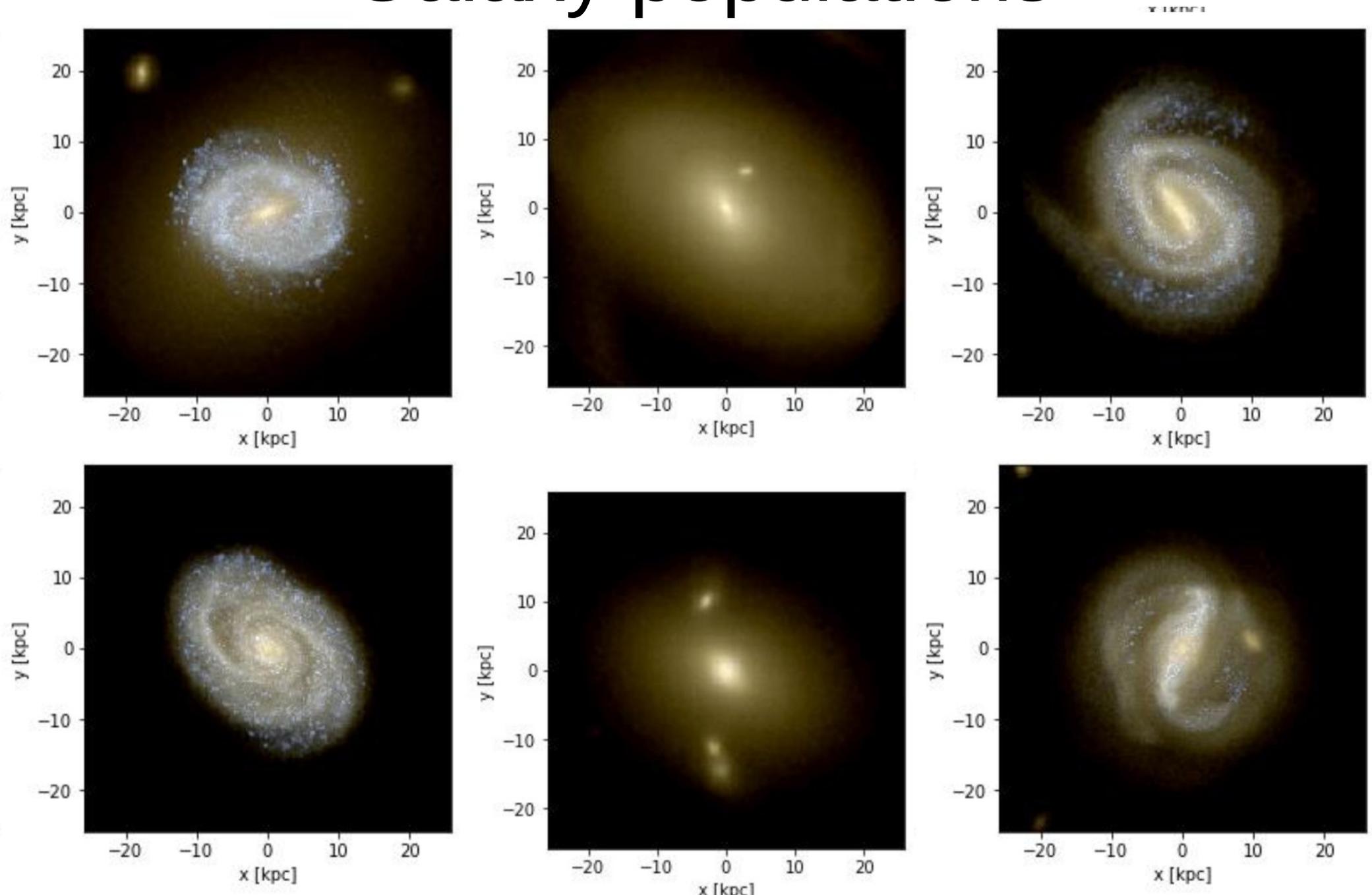
Galaxy Cluster



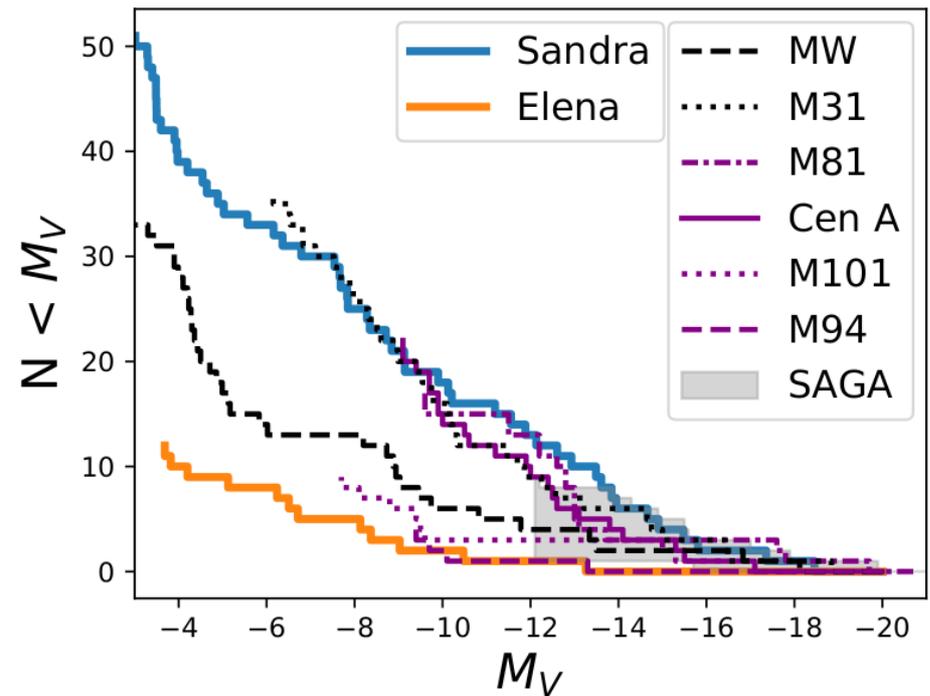
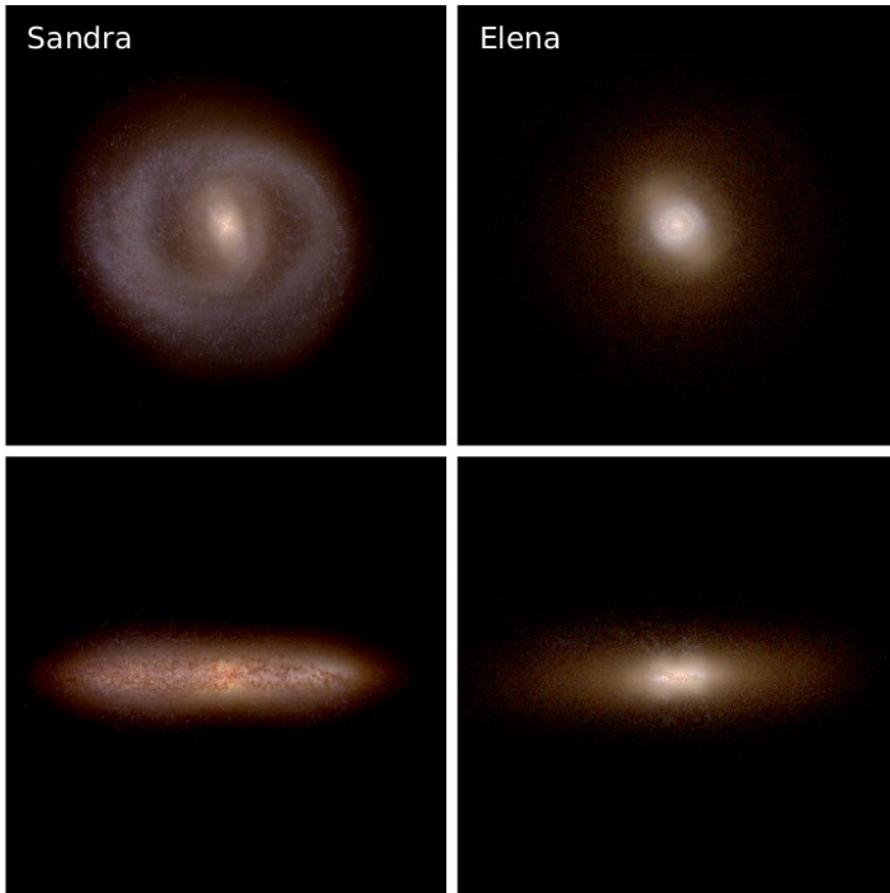
Galaxy Cluster Cores



Galaxy populations



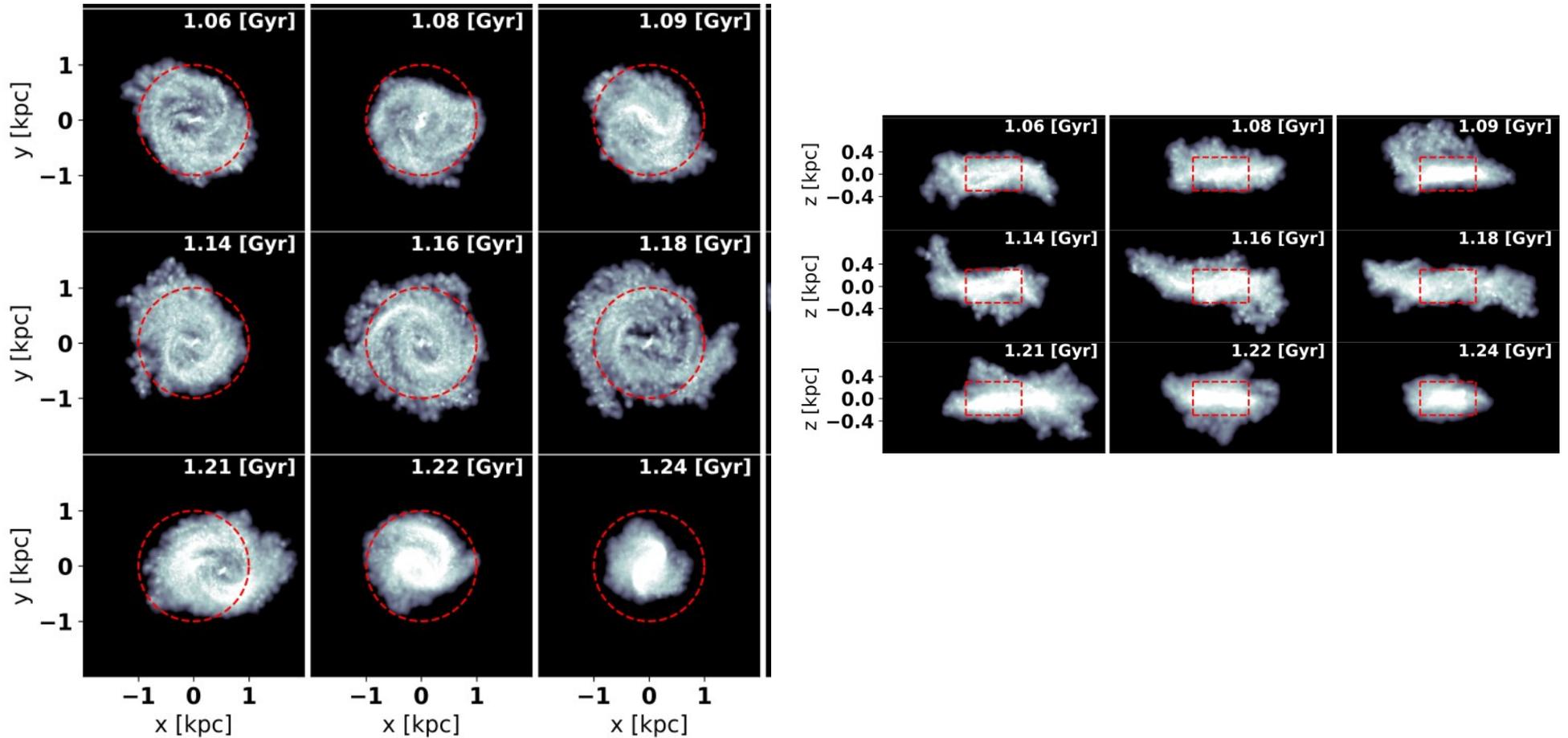
Faint Satellites of Galaxies



DC Justice League Galaxies running on Frontera

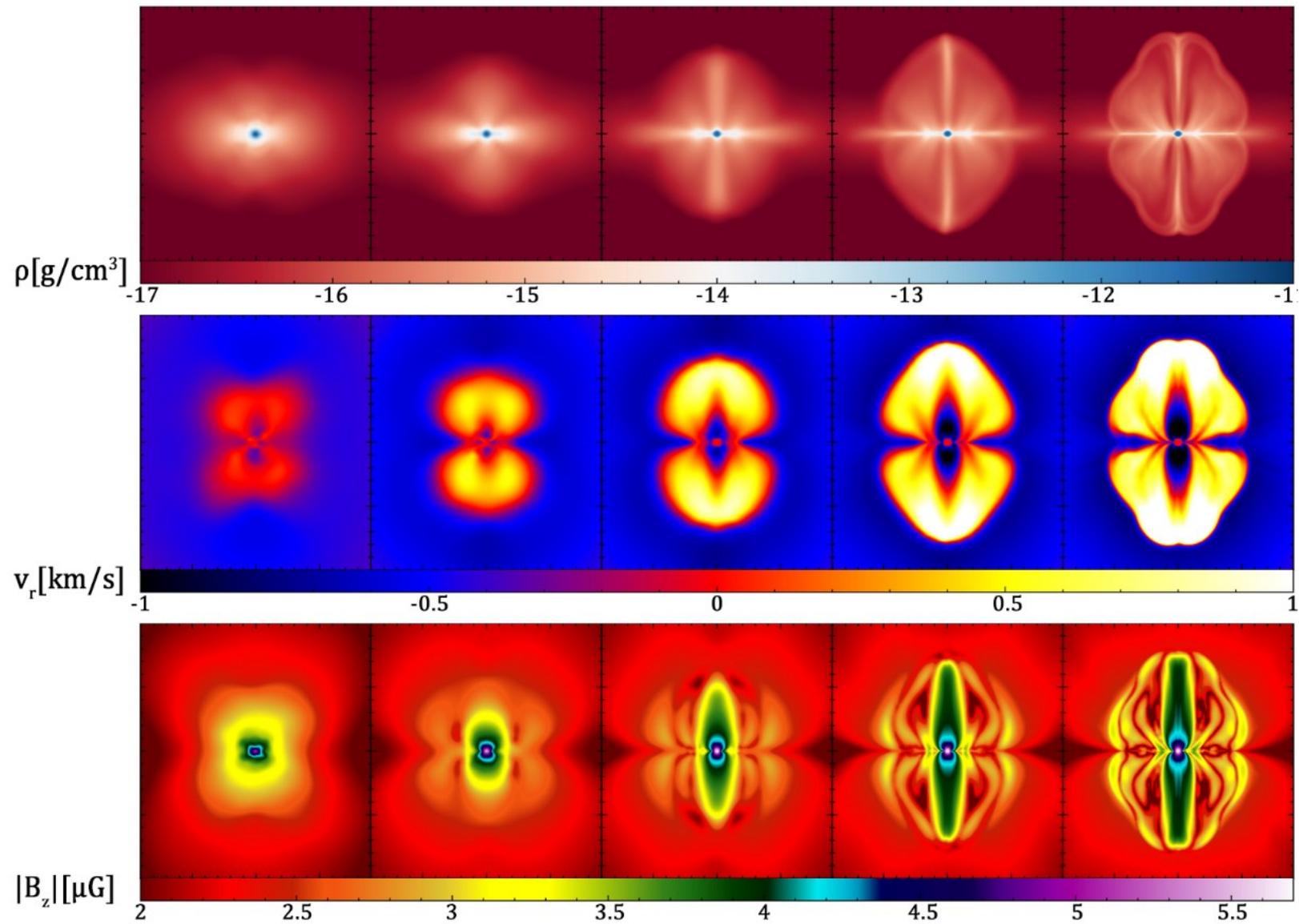
Applebaum et al 2021

Early Disk Formation



Galactic Disks form thin: Billion particle simulation on Piz Daint

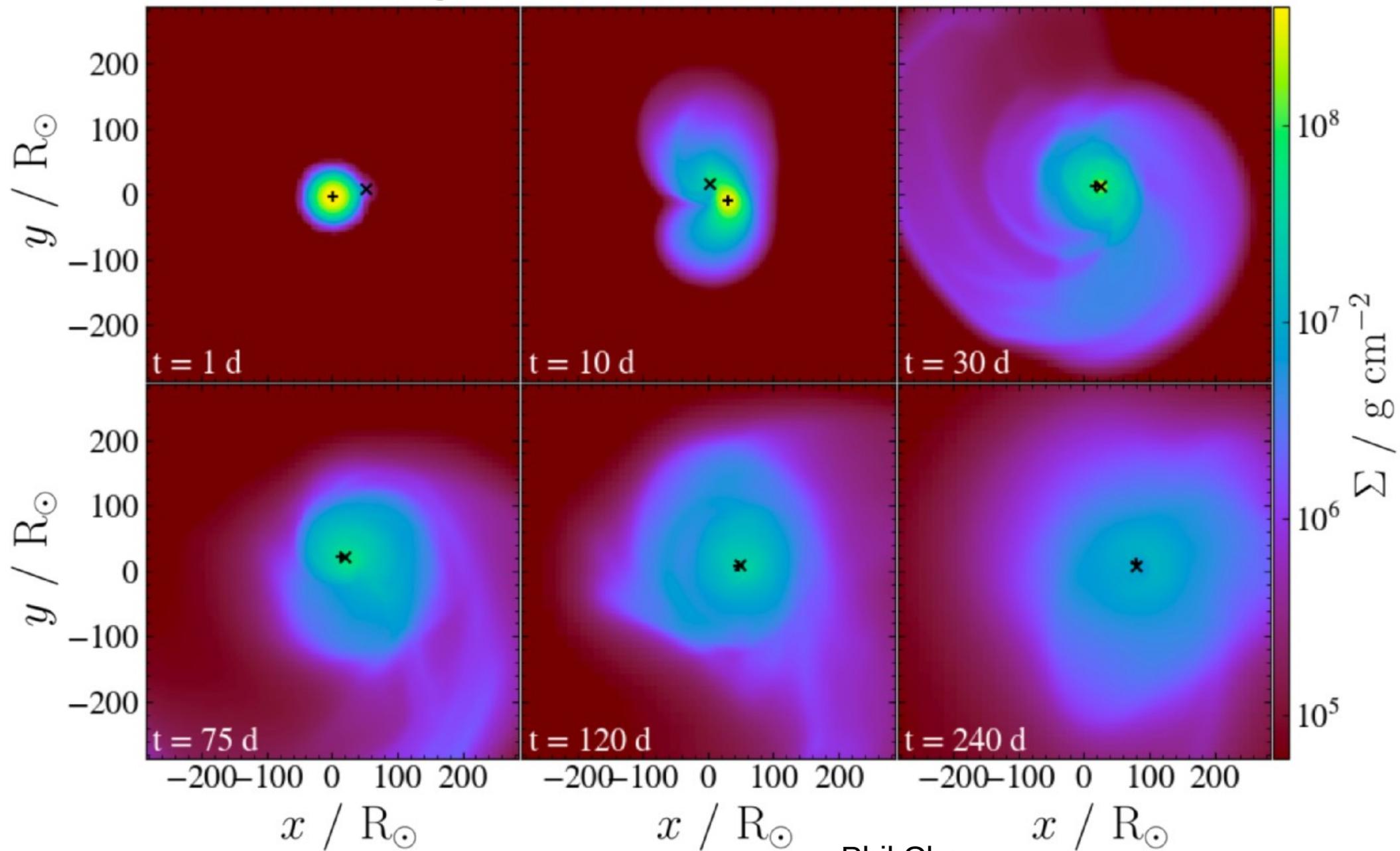
MHD Protostellar Cloud Collapse



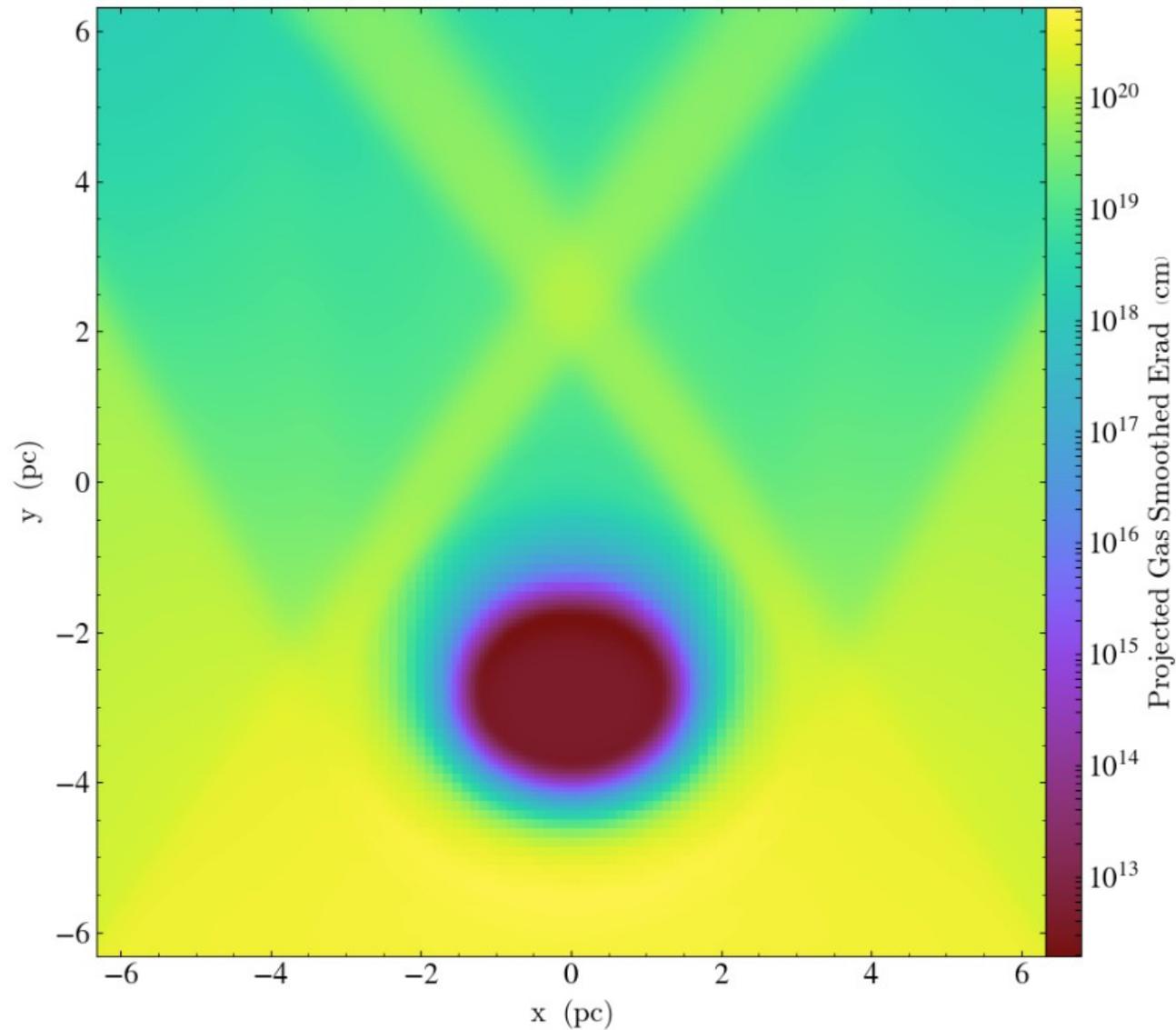
Moving Mesh Hydrodynamics

- More accurate hydrodynamics requires Riemann solvers
- Galilean invariance: mesh needs to follow the fluid flow
- Mesh needs to have arbitrary geometry
- Need a fast Voronoi mesh generator: ChaNGa (MANGA)

Binary Stars with MaNGA

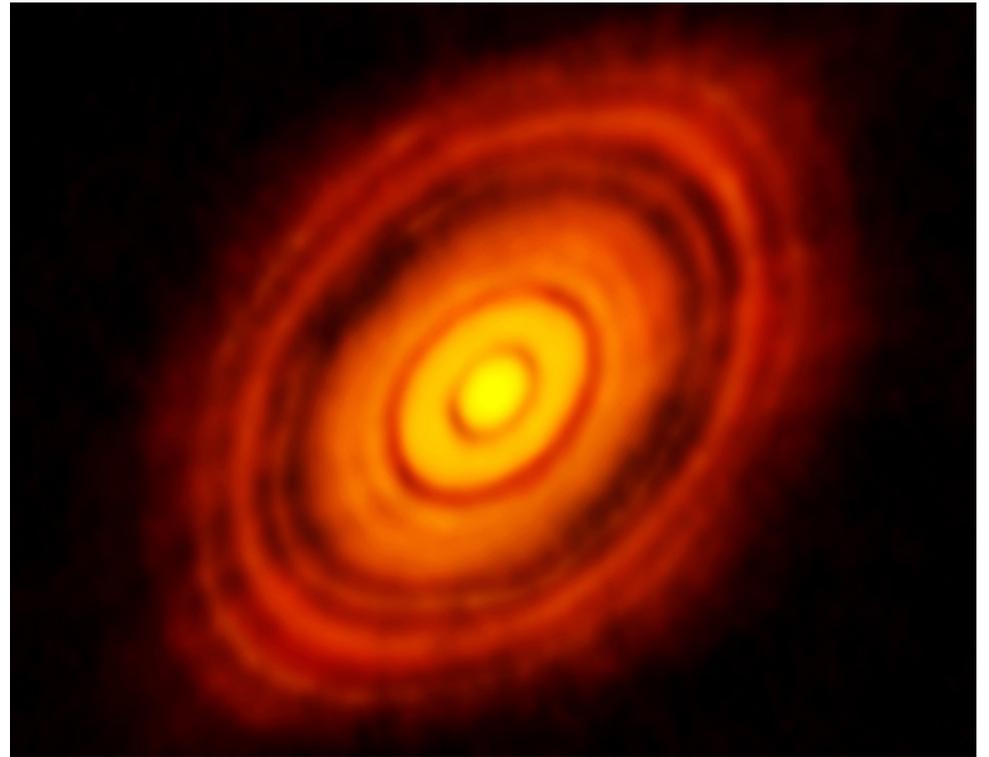


Radiation Hydrodynamics with MaNGa

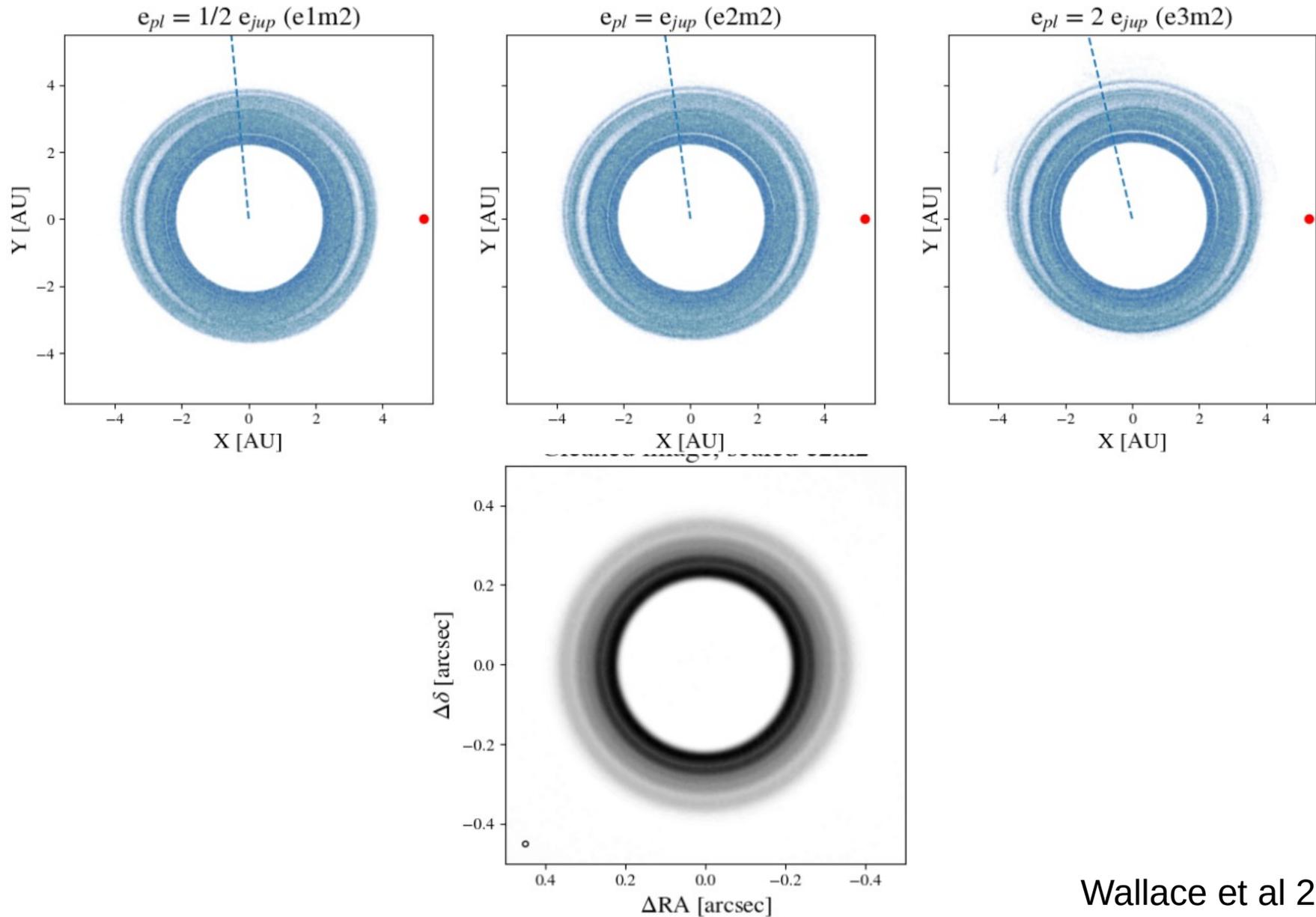


Protoplanetary Disks

- Likely result of cloud collapse with conserved angular momentum
- Disks can be gravitationally unstable
- Fragmentation depends on details of gas dynamics

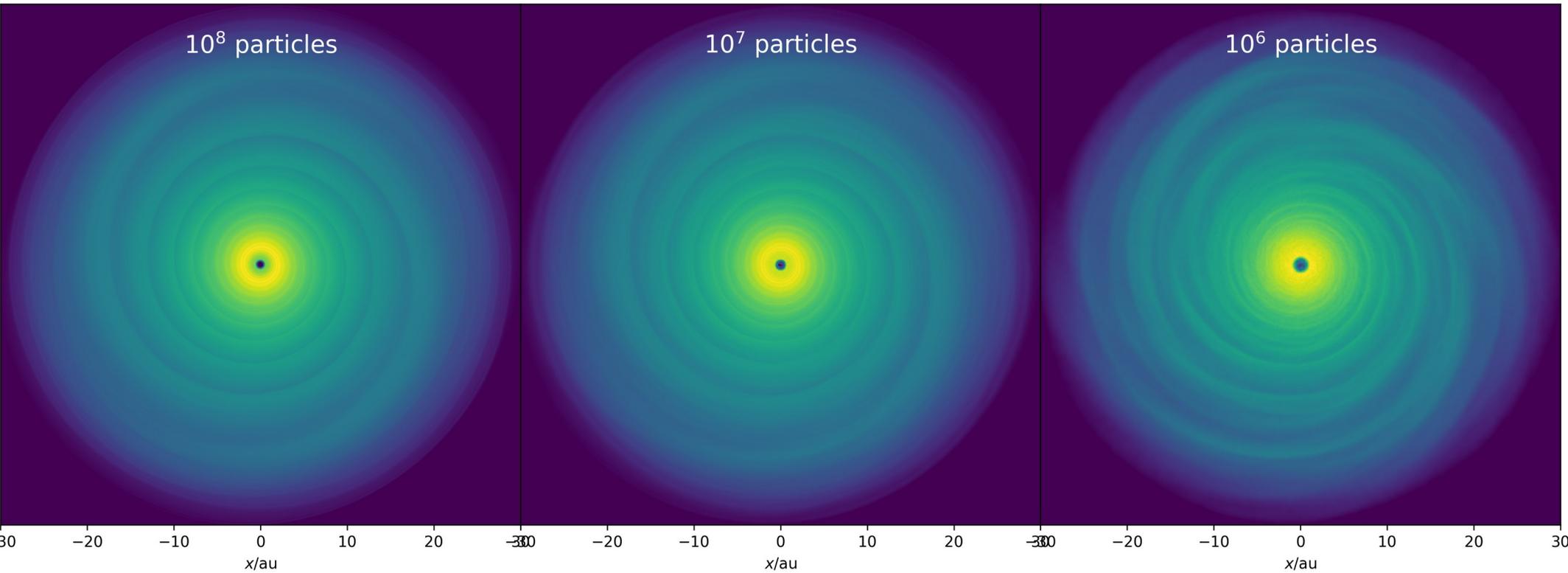


Dust from Collisions?



Resolution and Disks

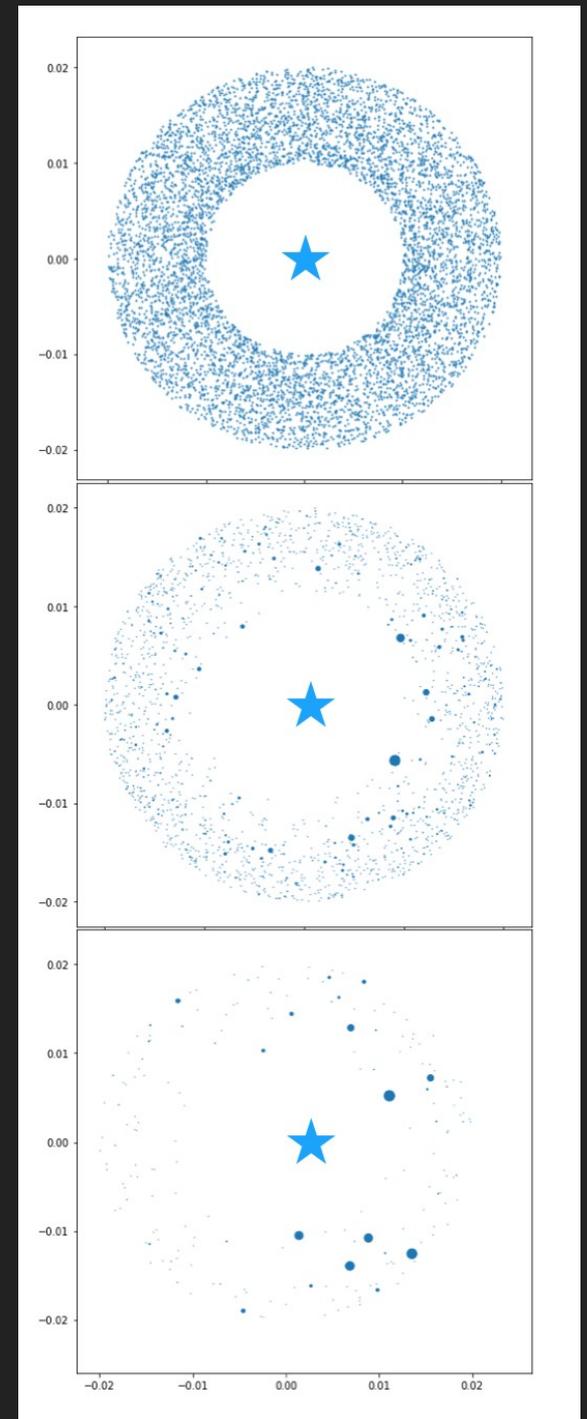
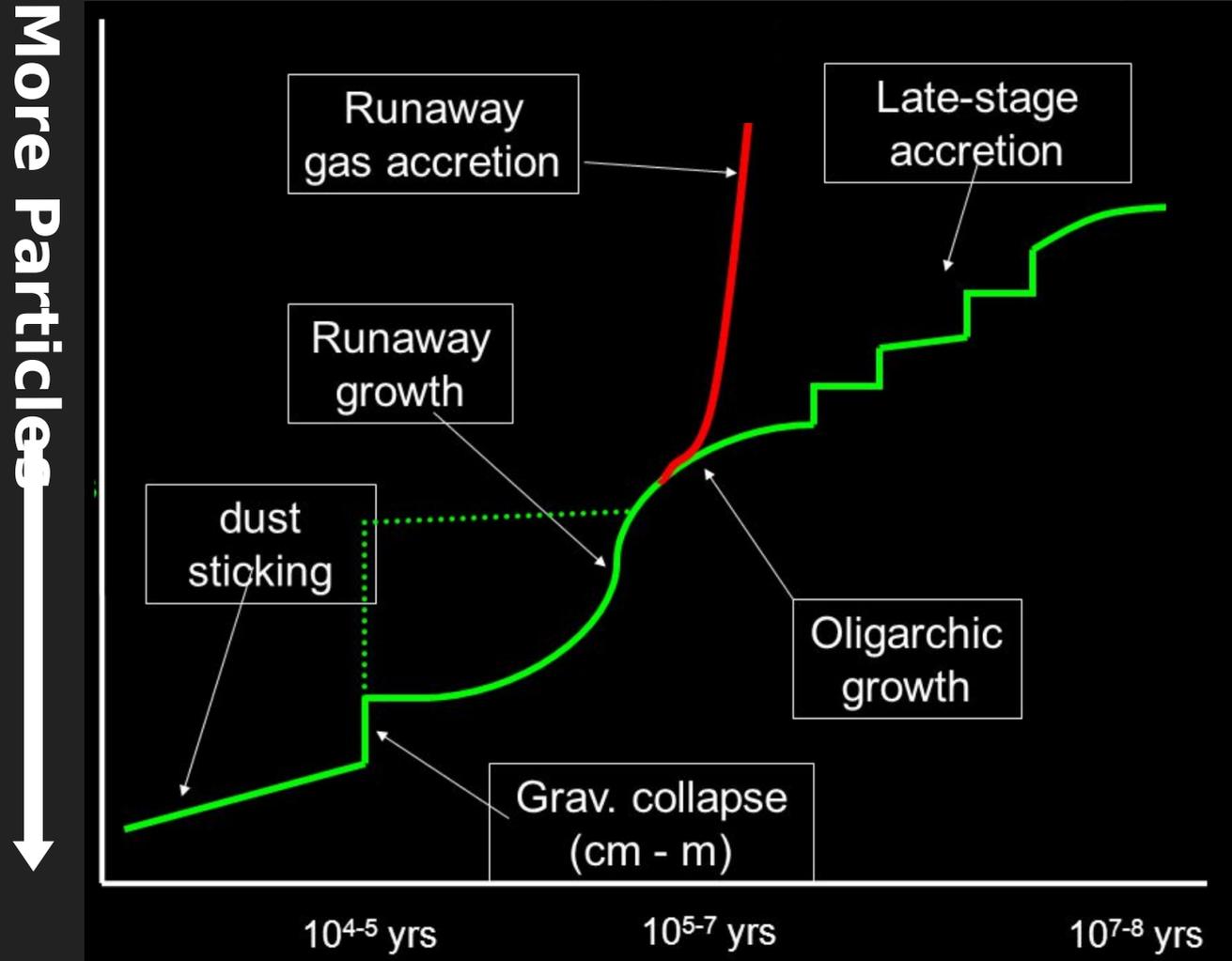
Resolution comparison: density
after 1.89 ORPs



Isaac Backus, Ph. D. Thesis

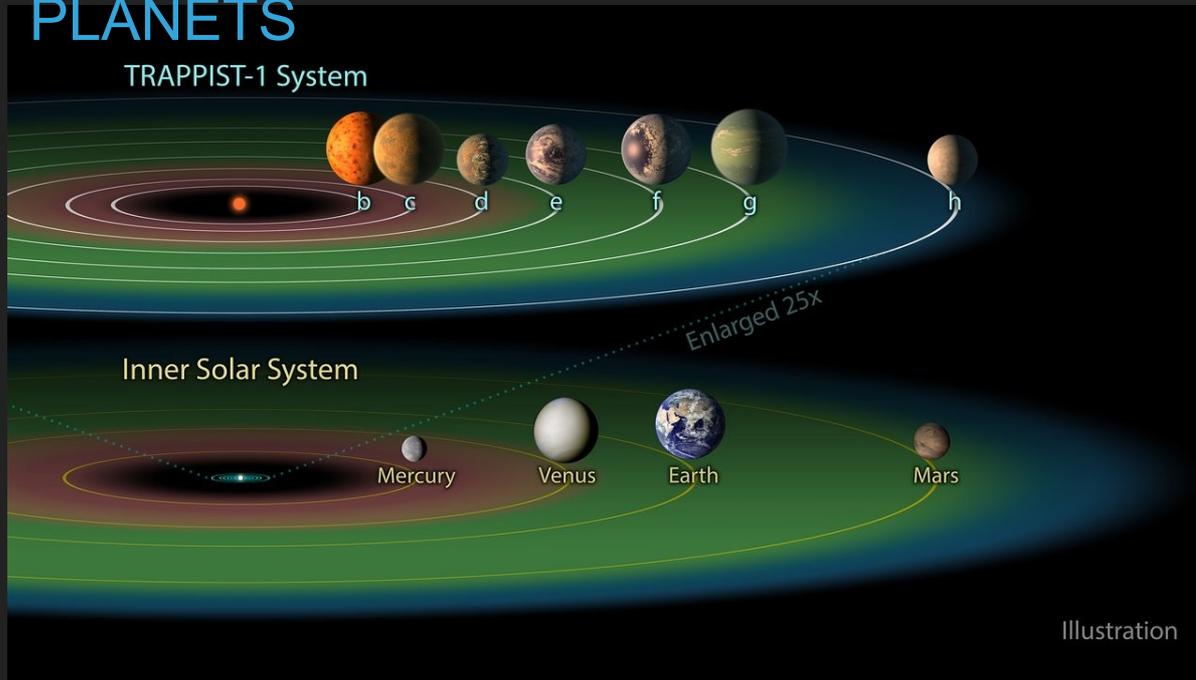
FROM DUST TO PLANETS

Image credit: Sean Raymond



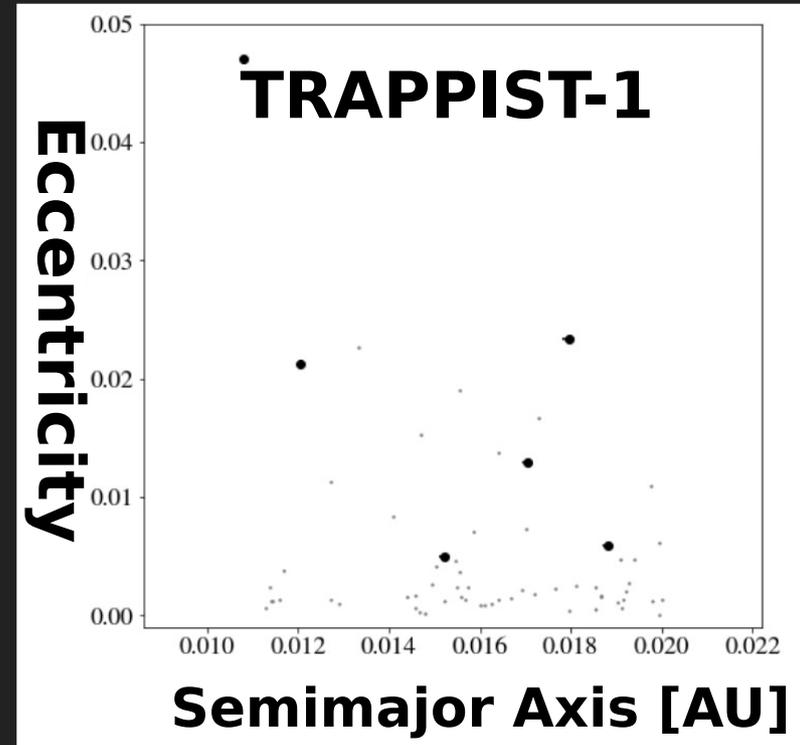
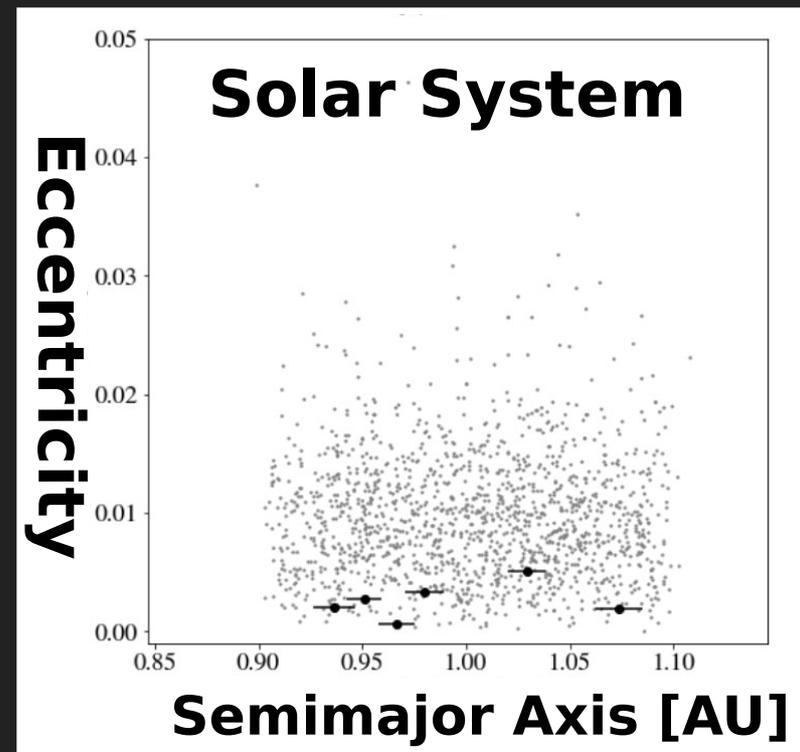
- ▶ ChaNGa can directly simulate accretion of planetesimals, test planetesimal formation models

SYSTEMS OF TIGHTLY PACKED INNER PLANETS

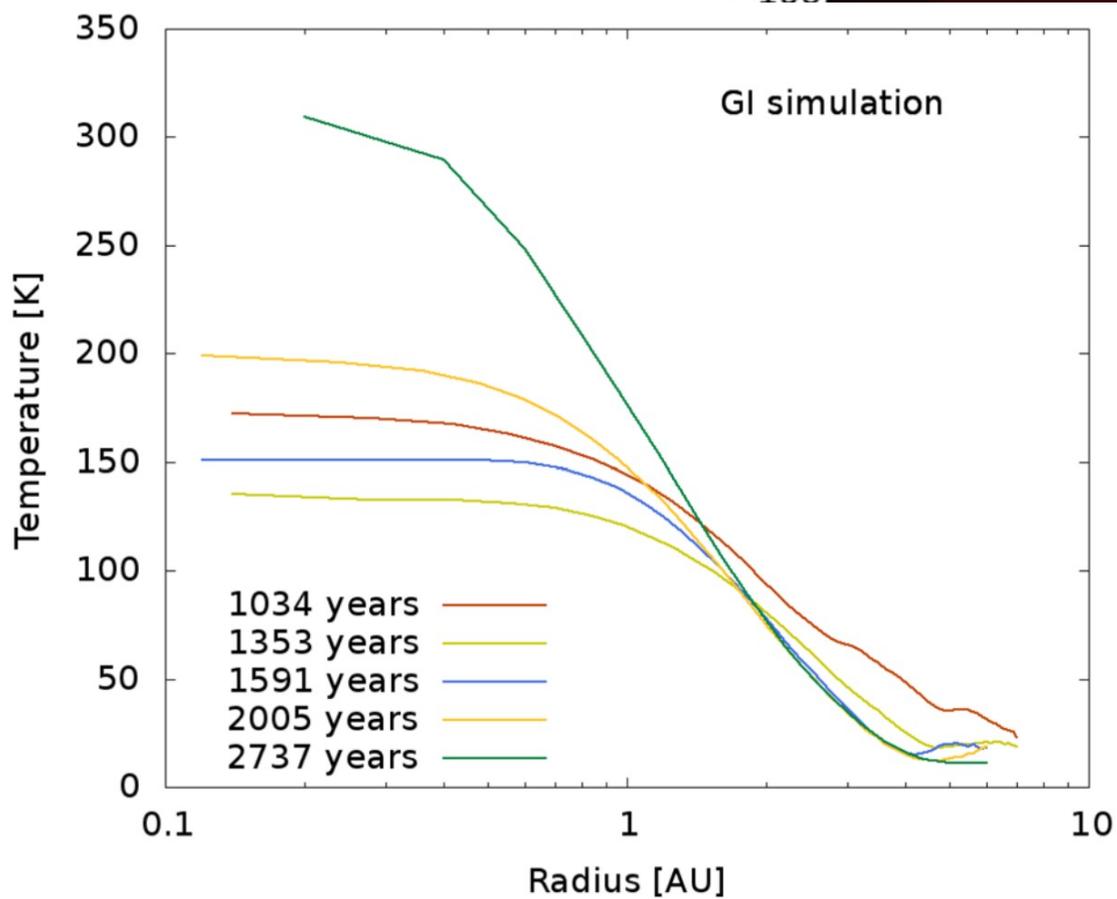
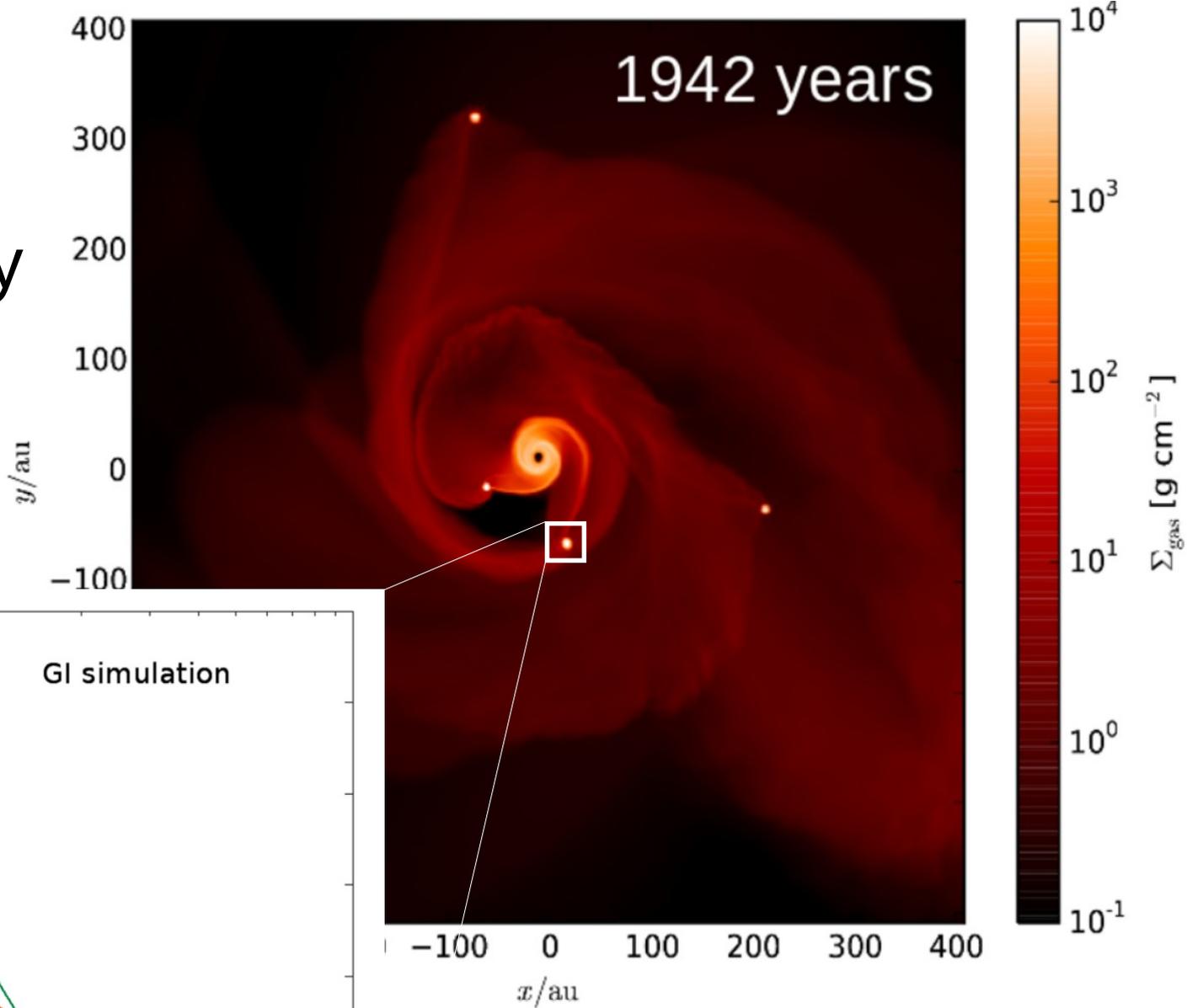


- ▶ High surface density, short dynamical timescale close to star
- ▶ Does the runaway + oligarchic growth model still apply?

Wallace, in prep.



Circum-Planetary Disks



Szulagy et al 2017

Summary

- Astrophysical simulations provide a challenges to parallel implementations
 - Non-local data dependencies
 - Hierarchical in space and time
- ChaNGa has been successful in addressing these challenges using Charm++ features
 - Computation/Communication overlap
 - Message priorities
 - Load Balancing
 - Modularity to add new Physics

Availability

- ChaNGa: <http://github.com/N-bodyShop/changa>
 - See the Wiki for a developer's guide
- Paratreet: <http://github.com/paratreet>
 - Some design discussion and sample code

User Community:

<https://nbody.shop>



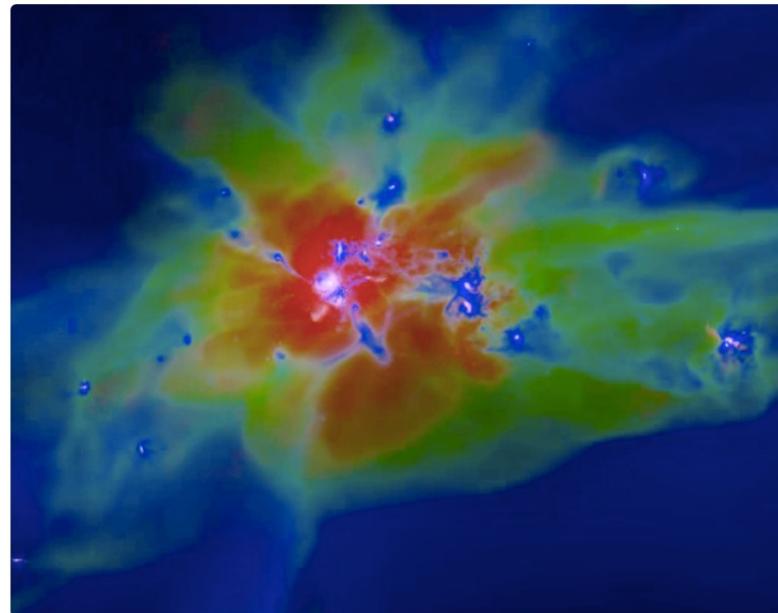
The N-Body Shop Conference



N-Body Shop Excellence Conference 2021

A VIRTUAL COLLABORATION CONFERENCE

[LEARN MORE](#)



Acknowledgments

- NSF ITR
- NSF Astronomy
- NSF SSI
- NSF XSEDE program for computing
- BlueWaters Petascale Computing
- Frontera Petascale Computing
- NASA HST
- NASA Advanced Supercomputing