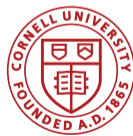


SpECTRE: Towards improved simulations of relativistic astrophysical systems

Nils Deppe

May 1, 2019



- ① Background and motivation
- ② Numerical methods
- ③ SpECTRE implementation

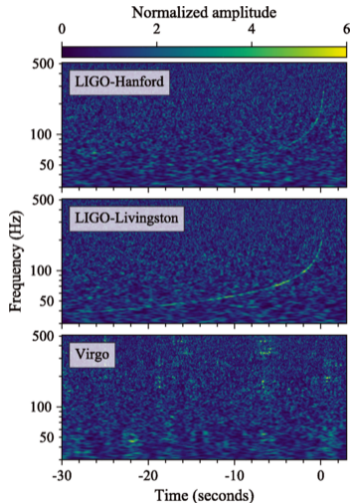
Simulations of GRMHD coupled to Einstein's equations are complicated, difficult, and interesting

- Accretion disks
- Binary neutron star mergers
- Core-collapse supernova explosions



Event Horizon Telescope Collaboration

- Gravitational waveforms for LIGO/Virgo and space-based detectors
- LIGO/Virgo follow-up waveforms
- Accretion for Event Horizon Telescope
- Improved understanding of heavy element generation



Abbott et al. 2017

- Hyperbolic equations in general form:

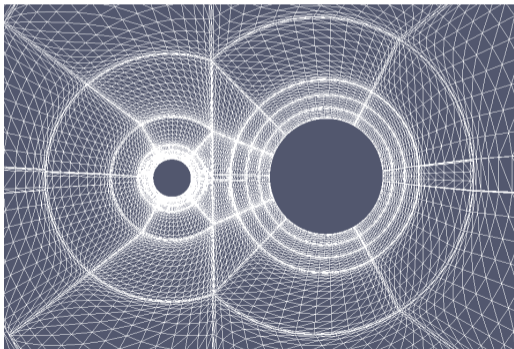
$$\partial_t \mathbf{U} + \partial_i \mathbf{F}^i(\mathbf{U}) + \mathbf{B}^i \cdot \partial_i \mathbf{U} = \mathbf{S}(\mathbf{U})$$

- Elliptic equations of the form:

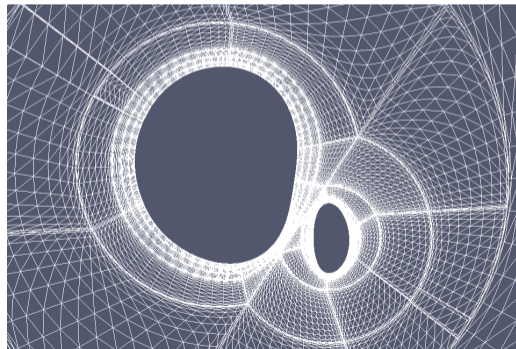
$$\partial^2 \mathbf{U} = \mathbf{S}(\mathbf{U}, \partial \mathbf{U})$$

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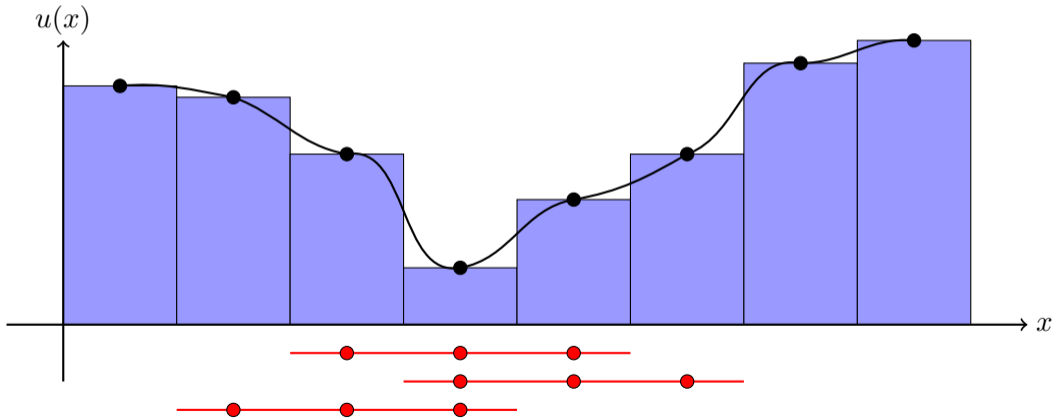
- Smooth solutions
- Exponential convergence



- Non-overlapping grids
- General grids:



- Work on shocks
- Polynomial convergence
- Typically Cartesian grids
- Overlapping grids

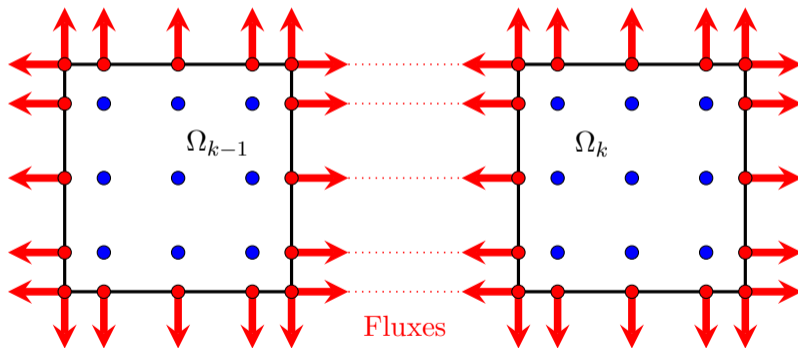


Current codes:

- Message passing (MPI) + some threading
- Spectral Einstein Code (SpEC):
 - Spectral methods: one element per core
 - Finite volume: $\sim 100,000 - 150,000$ cells per core
- Pseudospectral methods ~ 50 cores
- Finite volume methods $\sim 20,000$ cores

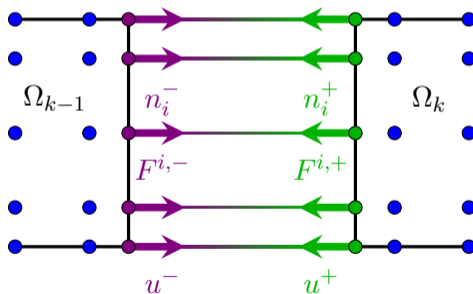
- Exponential convergence for smooth solutions
- Shock capturing
- Non-overlapping deformed grids
- *hp*-adaptivity
- Local time stepping
- Nearest-neighbor communication

- Boundary fluxes communicated between elements
- Nearest-neighbor only, good for parallelization



- Consider element Ω_{k-1} :

$$G_{k-1} = \frac{1}{2} (F^{i,+} n_i^+ + F^{i,-} n_i^-) - \frac{C}{2} (u^+ - u^-)$$



- 1 Compute time derivatives
- 2 Send data for boundary data
- 3 Integrate in time
- 4 Send data for limiting
- 5 Apply limiter

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- General framework for hyperbolic (Cornell, Caltech, CalState Fullerton, UNH) and elliptic (AEI) PDEs

- Scalar wave
- Curved scalar wave (mostly)
- Newtonian Euler (in code review)
- Relativistic Euler (mostly)
- GRMHD
- Generalized harmonic (in code review)

Numerical fluxes:

- Rusanov (local Lax-Friedrichs)
- HLL
- Upwind

Planned numerical fluxes:

- HLLC
- Roe
- Marquina

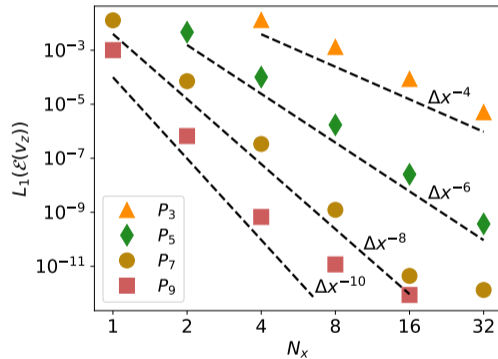
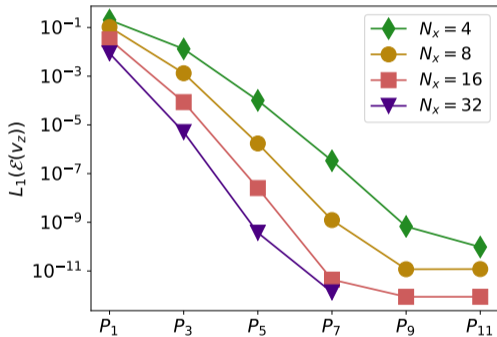
Limiters:

- Minmod (MUSCL, $\Lambda\Pi^1$, $\Lambda\Pi^N$)
- Krivodonova
- SimpleWENO (in code review)
- HWENO (in code review)
- Multipatch FV/FD subcell (in progress)

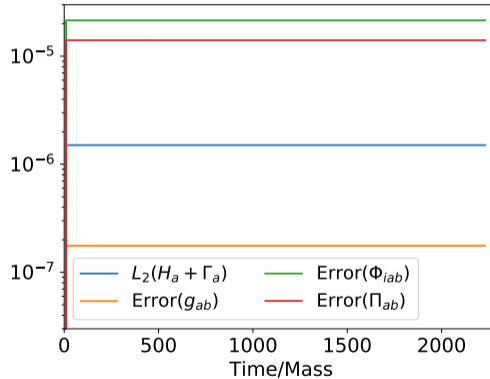
Planned limiters:

- Moe-Rossmanith-Seal (MRS)
- Hierarchical Barth-Jespersen and vertex-based

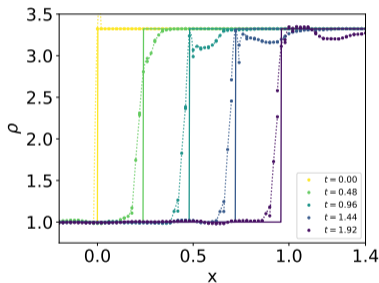
Convergence for Smooth Problems: Alfvén Wave



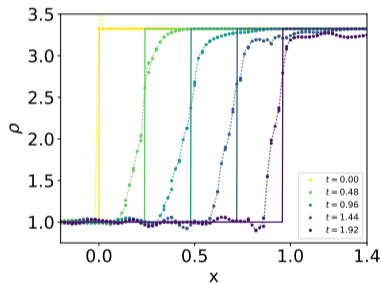
- Generalized harmonic system
- Excised cube in center



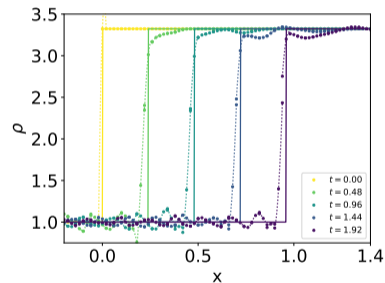
$256 \times 1 \times 1$ elements, 3^3 points per element



Krivodonova

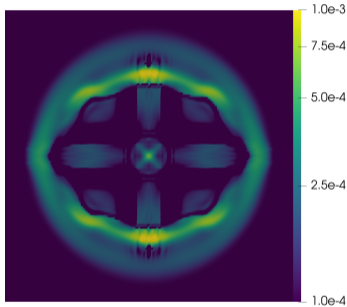


SimpleWENO

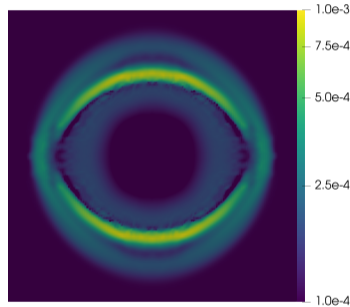


HWENO

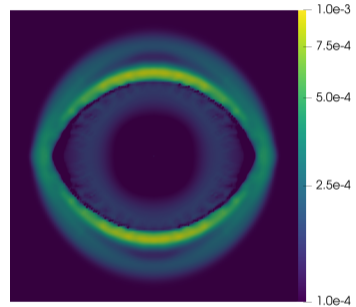
$128^2 \times 1$ elements, 2^3 points per element



Krivodonova

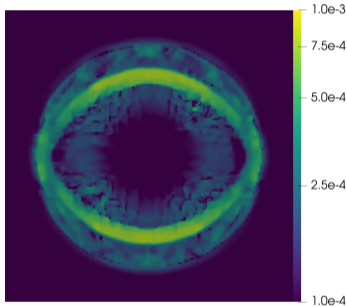


SimpleWENO

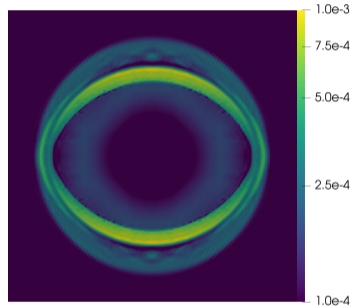


HWENO

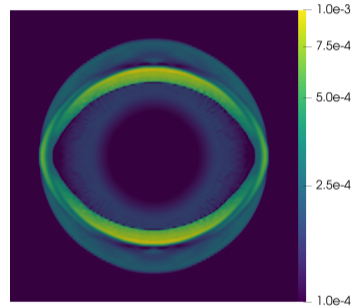
$128^2 \times 1$ elements, 3^3 points per element



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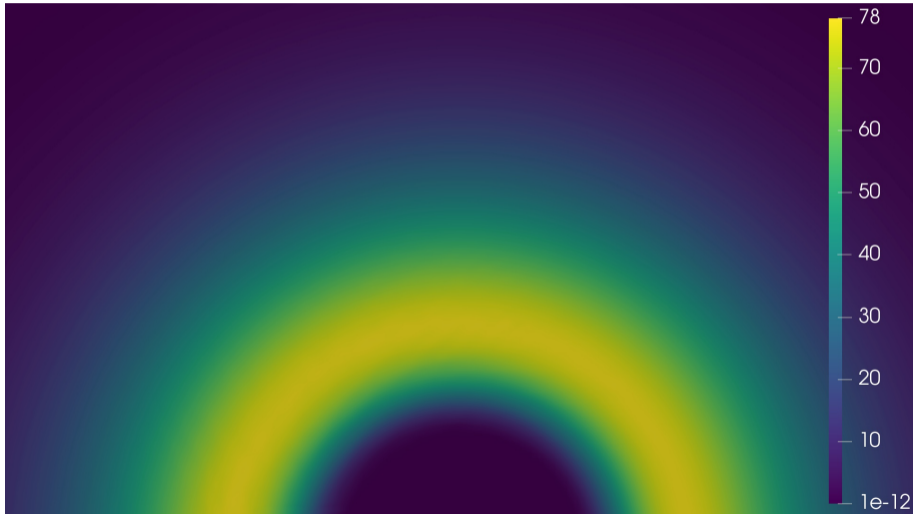
SimpleWENO



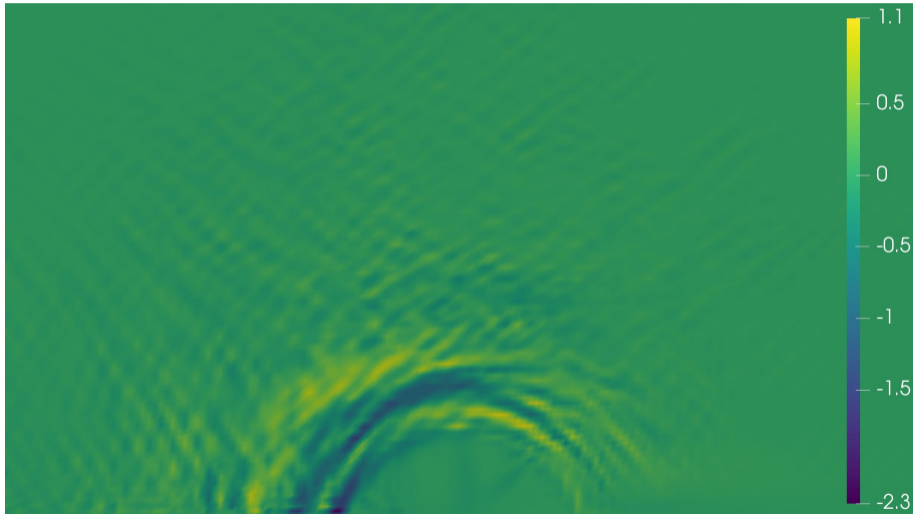
HWENO

- Torus around a black hole
- Code comparison project
- $\chi = 0.9375$, $\rho_{\max} \approx 77$
- Orbital period $T_{\text{orb}} \approx 247$
- Hexahedron: $[-40, 40] \times [2, 40] \times [-8, 8]$

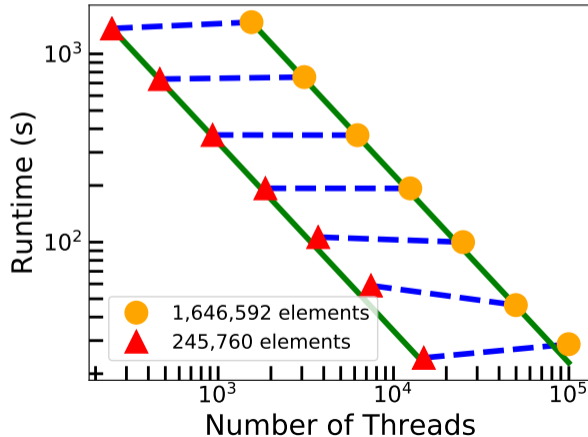
Rest mass density ρ at $t = 600$



Error in rest mass density ρ at $t = 600$



- Run on BlueWaters supercomputer, NCSA, UIUC, IL, USA
- Green is perfect speedup for fixed problem size (strong scaling)
- Blue shows actual weak scaling (flat is ideal)



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- Limiting and primitive recovery an open problem